

Understanding Neutrino Properties, and their Mass Measurement

Particle Physics Seminar, University of Birmingham, Nov. 3, 2021

Christian Weinheimer – Institute for Nuclear Physics, University of Münster

- Current key issues in neutrino physics
- Highlights from
 - neutrino oscillation experiments
 - search for $0\nu\beta\beta$
- Direct search for $m(\nu)$
 - recent results from KATRIN
- Conclusions



Neutrinos in the Standard Model

- extended by neutrino masses -

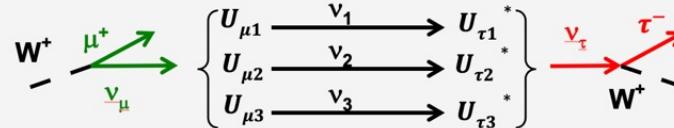
Neutral, spin $\frac{1}{2}$, 3 flavours, only weak interaction

3 flavour eigenstates and mass eigenstates differ:

$m(\nu) \neq 0$ and ν mixing

$$U_{PMNS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \cdot \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{i\delta_{CP}} & 0 & \cos \theta_{13} \end{pmatrix} \cdot \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \cdot \begin{pmatrix} e^{i\eta_1} & 0 & 0 \\ 0 & e^{i\eta_2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

→ ν flavour oscillation, CP violation via U_{PMNS} possible

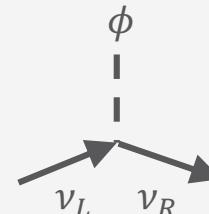


Neutral: particle character unclear:

Dirac or Majorana ($\nu = \bar{\nu}$) ?

Neutrino masses so much lighter:

Yukawa coupling to the Higgs ?



Standard Model of Elementary Particles		
three generations of matter (fermions)		
QUARKS		interactions / force carriers (bosons)
mass charge spin	mass charge spin	mass charge spin
$=2.2 \text{ MeV}/c^2$ $\frac{2}{3}$ $\frac{1}{2}$ u	$=1.28 \text{ GeV}/c^2$ $\frac{2}{3}$ $\frac{1}{2}$ c	$=173.1 \text{ GeV}/c^2$ $\frac{2}{3}$ $\frac{1}{2}$ t
down	strange	bottom
$=4.7 \text{ MeV}/c^2$ $-\frac{1}{3}$ $\frac{1}{2}$ d	$=96 \text{ MeV}/c^2$ $-\frac{1}{3}$ $\frac{1}{2}$ s	$=4.18 \text{ GeV}/c^2$ $-\frac{1}{3}$ $\frac{1}{2}$ b
electron neutrino	muon neutrino	tau neutrino
$<1.0 \text{ eV}/c^2$ 0 $\frac{1}{2}$ e	$<0.17 \text{ MeV}/c^2$ 0 $\frac{1}{2}$ muon neutrino	$<18.2 \text{ MeV}/c^2$ 0 $\frac{1}{2}$ tau neutrino
electron	muon	tau
$=0.511 \text{ MeV}/c^2$ -1 $\frac{1}{2}$ e	$=105.66 \text{ MeV}/c^2$ -1 $\frac{1}{2}$ muon	$=1.7768 \text{ GeV}/c^2$ -1 $\frac{1}{2}$ tau
Z boson		W boson
$=91.19 \text{ GeV}/c^2$ 0 1 Z		$=80.39 \text{ GeV}/c^2$ ± 1 1 W
SCALAR BOSONS		
GAUGE BOSONS VECTOR BOSONS		

modified from en.wikipedia.org

Current key issues

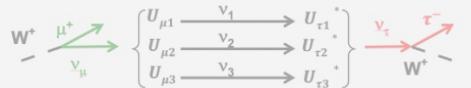
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→ ν flavour oscillation, CP violation via U_{PMNS} possible

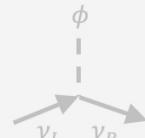


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SM BSM

1. Ordering: $m(\nu_3) > m(\nu_{2,1})$ (NO) or $m(\nu_{2,1}) > m(\nu_3)$ (IO) ?

2. CP violating phase $0 \neq \delta_{CP} \neq \pi$?

3. Lepton-number violation by Majorona ν ?

4. Absolute neutrino mass

neutrino mass generation likely BSM

10^9 more ν than atoms → cosmological relevant: structure formation

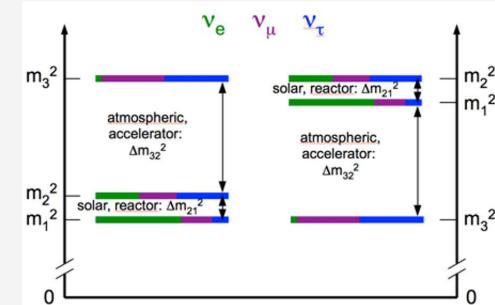
5. Is there a 4th or even a 5th light sterile neutrino ?

6. Are there more BSM effects ?

BSM via coherent elastic neutrino nucleon scattering CEvNS ?

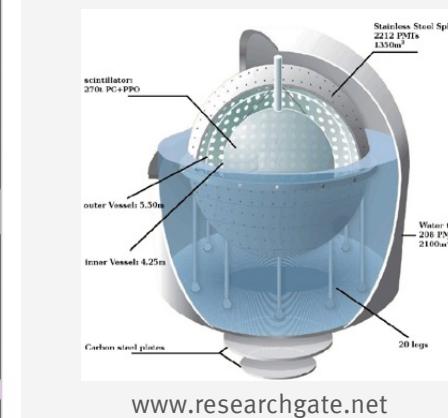
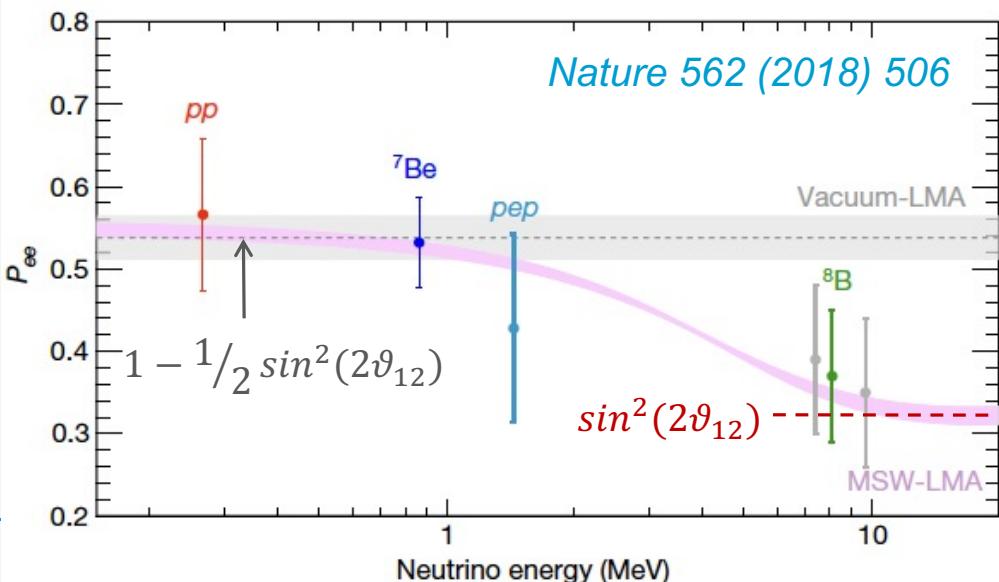
In addition, ν are interesting messenger for other physics:

e.g. from fusion inside the sun, from supernovae or UHE neutrino from “cosmic accelerators”

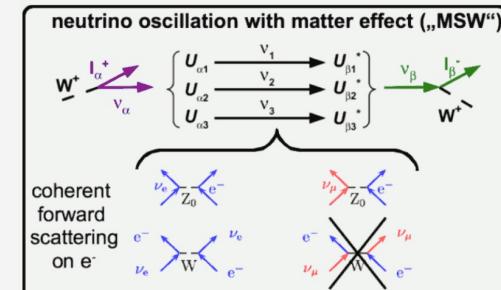
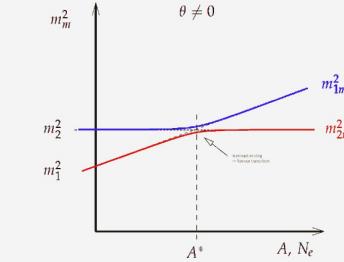
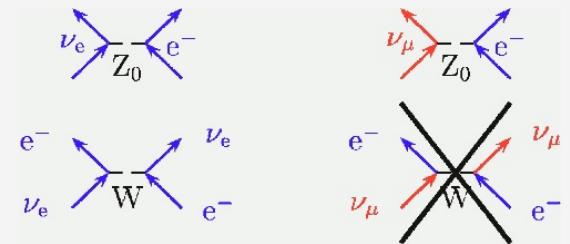


Borexino proved matter-enhanced oscillation (MSW)

- Matter-enhanced neutrino oscillation is due to the additional CC coherent forward scattering of ν_e on e^- in matter in contrast to ν_μ, ν_τ
- MSW: nearly complete ν flavour transformation possible**
if the electron density meets the resonance condition, e.g. for ^8B solar ν_e
& if the neutrino mass ordering is matching, here $m(\nu_2) > m(\nu_1)$



solar ν: $m(\nu_2) > m(\nu_1)$



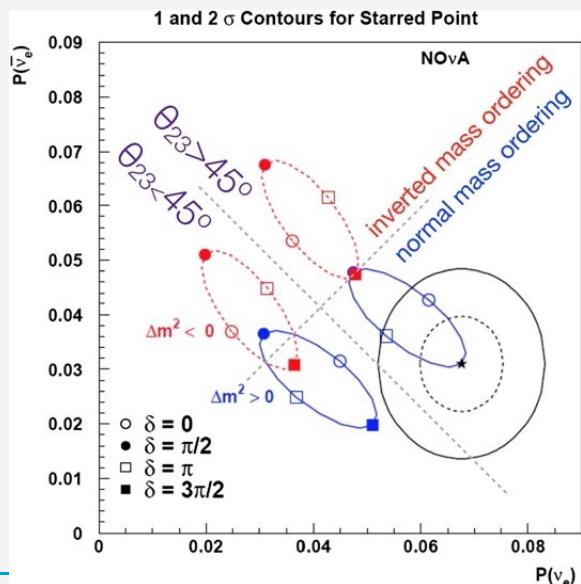
Long baseline: ordering and CP violation

ν oscillation in vacuum: $P(\nu_\alpha \rightarrow \nu_\beta) = \left| \sum_j U_{\alpha j}^* U_{\beta j} e^{-i \frac{m_j^2}{2E} L} \right|^2 = \delta_{\alpha\beta} - 4 \sum_{i>j} \Re(U_{\alpha i}^* U_{\alpha j} U_{\beta i} U_{\beta j}^*) \sin^2 \left(\frac{\Delta m_{ij}^2}{4E} L \right) + 2 \sum_{i>j} \Im(U_{\alpha i}^* U_{\alpha j} U_{\beta i} U_{\beta j}^*) \sin^2 \left(\frac{\Delta m_{ij}^2}{2E} L \right)$

Look for ν_e appearance in ν_μ off-axis beam: $P(\nu_\mu \rightarrow \nu_e)$, compare to $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$

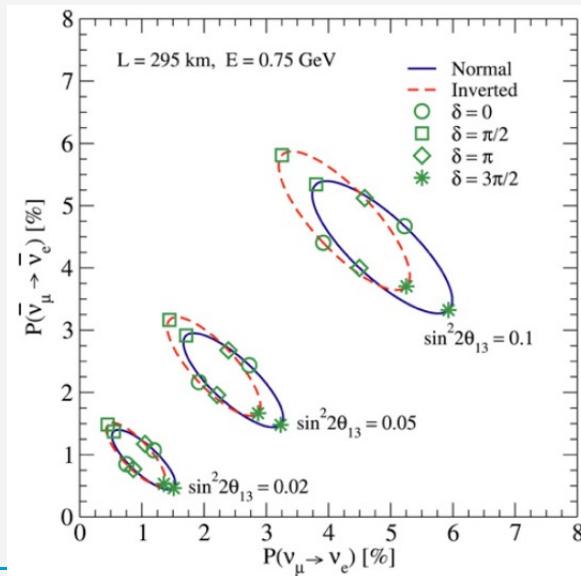
with CP-violation (CP-phase $\delta \neq 0$ or π) and with matter effects (L-dependent):

$$P(\nu_\mu \rightarrow \nu_e) \neq P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \quad (\rightarrow \text{still ellipses in bi-event plot})$$

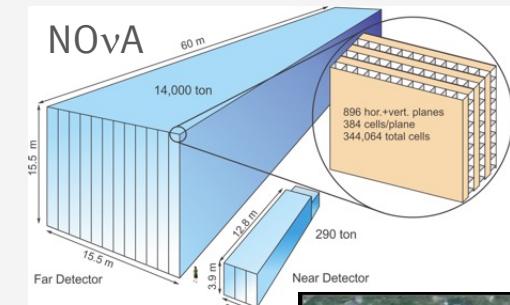
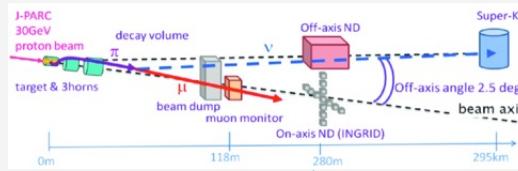
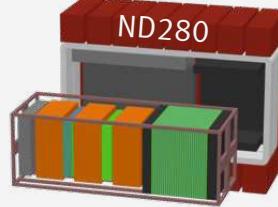
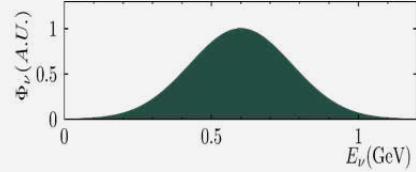
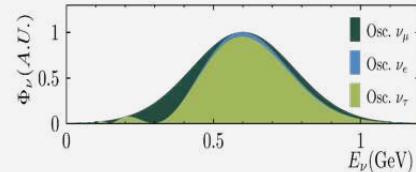


NOvA, 810 km
M.D. Messier,
Nucl. Phys. B 908 (2016) 151

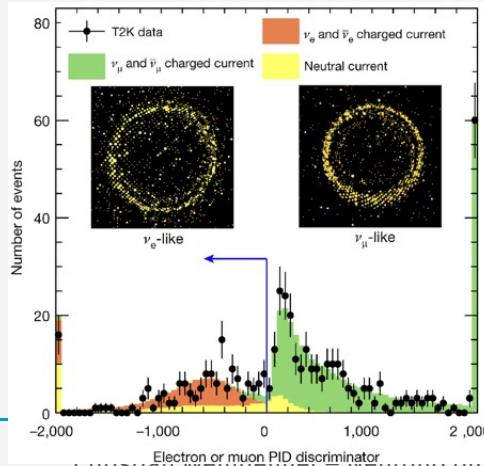
T2K
H. Nunokawa, S. Park,
J.W.F. Valle
Prog. Part. Nucl. Phys.
60 (2008) 338



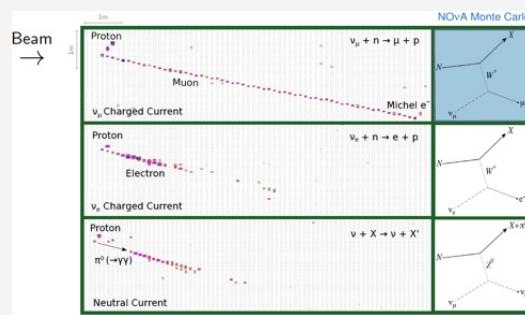
NOvA, T2K: Search for $\nu_\mu \rightarrow \nu_e, \nu_\mu$ and $\bar{\nu}_\mu \rightarrow \bar{\nu}_e, \bar{\nu}_\mu$



M. Strait,
Neutrino
Telescopes
2021



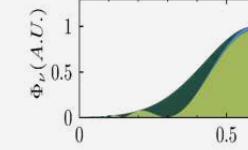
Christian Weinheimer – Neutrino properties and mass – Seminar Birmingham, Nov. 2021



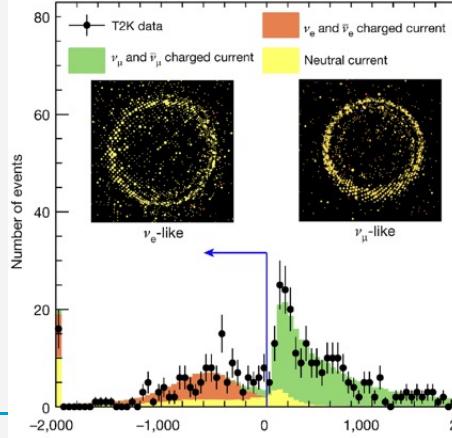
NOvA, T2K: Search for $\nu_\mu \rightarrow \nu_\mu$ and $\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$



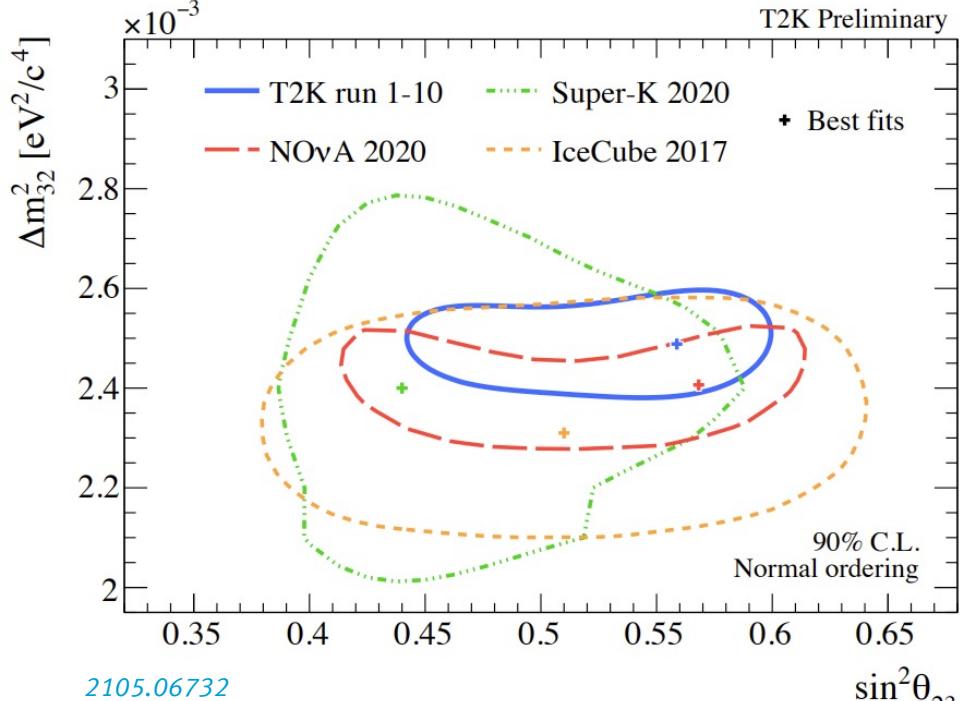
Super K

 Mt. Ikeno-Yama
1360 m

courtesy

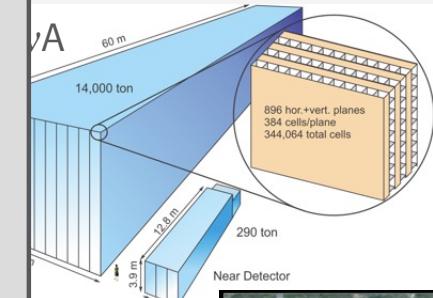
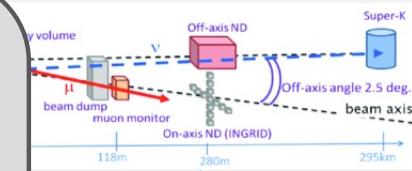
C. D. L.



$$N_\mu(E_\nu) = P(\nu_\mu \rightarrow \nu_\mu) \cdot \sigma(E_\nu) \cdot \Phi(E_\nu) \cdot \epsilon(E_\nu)$$

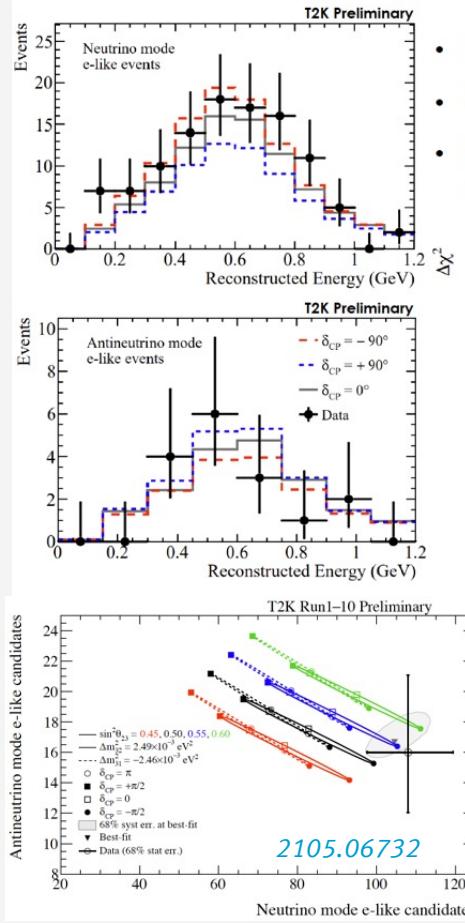


$\sin^2 \theta_{23}$: degeneracy w.r.t. octant
good agreement among all experiments



M. Strait,
Neutrino
Telescopes
2021

NOvA, T2K: Search for $\nu_\mu \rightarrow \nu_e$ & $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ appearance

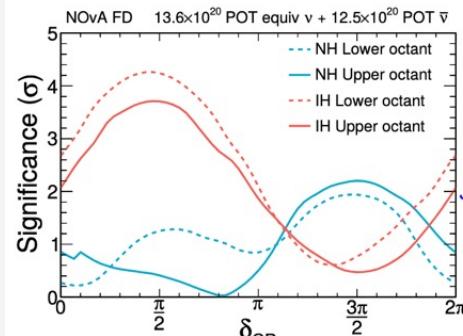


- Large region excluded at 3σ
- CP-conservation ($0, \pm$) excluded at 90%
- Weak preference for normal mass ordering

courtesy S. Dolan

$$N_e(E_\nu) = P(\nu_\mu \rightarrow \nu_e) \cdot \sigma(E_\nu) \cdot \Phi(E_\nu) \cdot \epsilon(E_\nu)$$

M. Strait,
Neutrino Telescopes 2021



Please note different offset

T2K: runs 1–10, 33% more protons for ν_μ run w.r.t. *Nature* 580 (2020) 339
excludes CP conservation at 90% C.L.

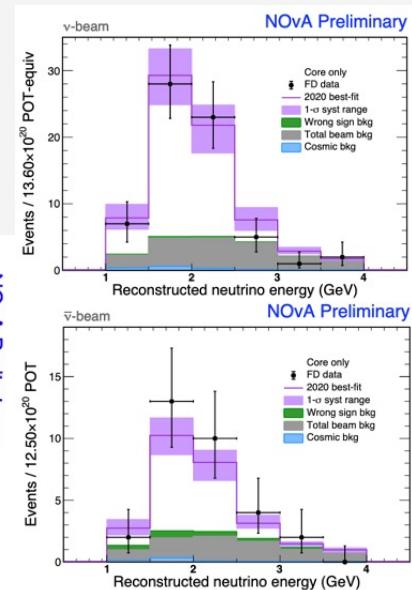
NOvA and T2K: both prefer NO, IO: δ_{CP} in good agreement, NO: some tension w.r.t. δ_{CP}

Discrepancy is statistically not compelling yet !

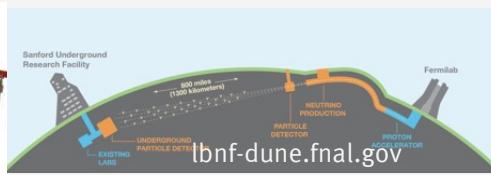
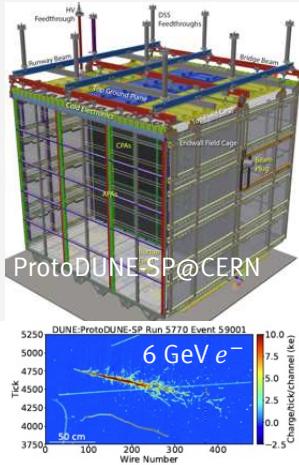
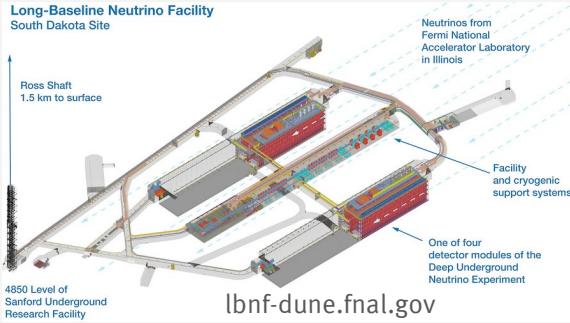
NOvA and T2K are performing common analysis

NOvA and T2K both are going to increase statistics by factor 2.5

& improve systematics (e.g. upgrade of beam power, ND280, Gd in SK)



Determination of δ_{CP} : DUNE and Hyper Kamiokande

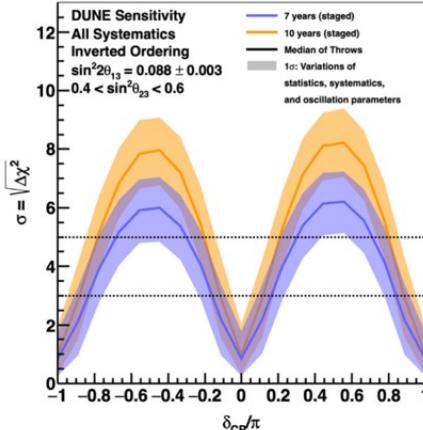


detectors 1,2: each 17 kt LAr
detector 1: vertical drift

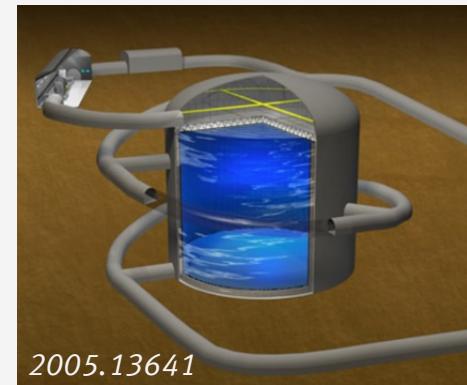
detector 2: probably horizontal drift

evacuation started, ν beam in 2029

later **THEIA?** (water-based liquid scintillation detector: WbLS)



Hyper-Kamiokande: 260 kt water
→ 188kt fiducial (8.4x SK)



increase beam power: 500 kW → 1.3 MW

increase horn current: 250 kA → 320 kA

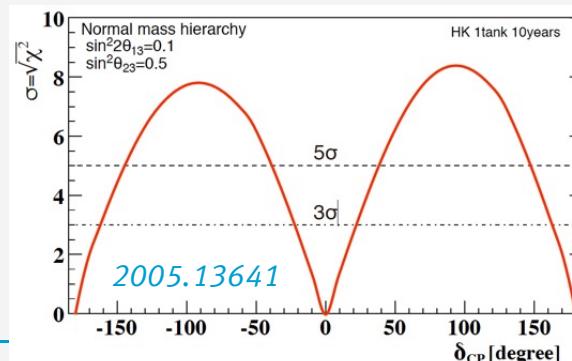
→ 10% more neutrinos & less wrong sign

upgrade near detectors, e.g. ND280

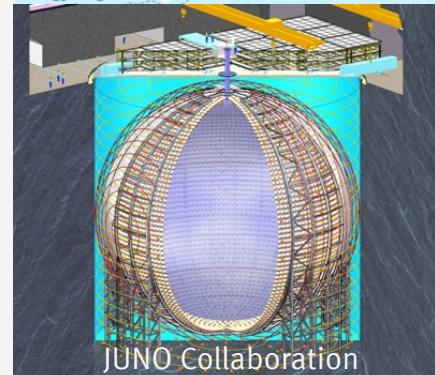
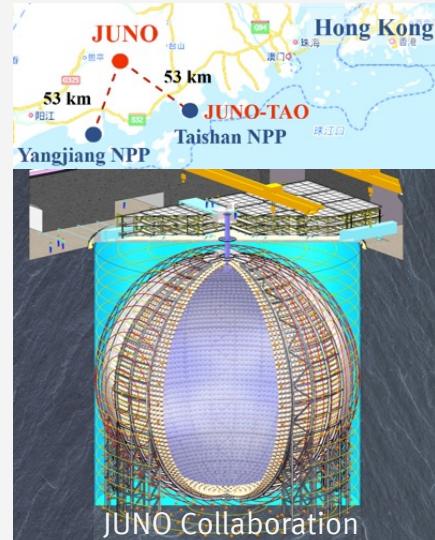
P. de Perio, CAP 2021

Excavation started
begin water filling & commissioning in 2027

Other physics:
solar ν , nucleon decay,
supernova ν , ...



JUNO, 20 kT liquid scintillator



Oscillation physics:

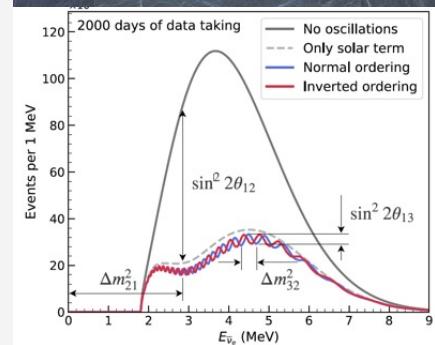
- distinguish mass ordering by 3σ after 6 yrs of exposure, better together with LBL
- improvement of $\sin^2 2\theta_{12}$, Δm_{21}^2 , $|\Delta m_{32}^2|$ to 0.6% after 6 yrs (2104.02565)
- 2 planned Taishan reactor cores will not be built in the near term → less statistics but compensated nearly by gain in photo-efficiency, linearity and other factors

Non-oscillation neutrino physics:

- solar neutrinos: ${}^8\text{B}$ 60k events at 30k bg, helps differentiate low/high metallicity models
- geo neutrinos: down to 5% flux precision
- supernova neutrinos: a few 1000 events (ν_e and $\nu_{\mu,\tau}$) for galactic core-collapse SN
- Diffuse supernova background: 3σ after 10 yrs
- p-decay: expect to improve w.r.t. Super Kamiokande in kaon and possibly other channels

Timeline:

- tunnel und underground hall excavation finished, expected to complete construction end of 2022



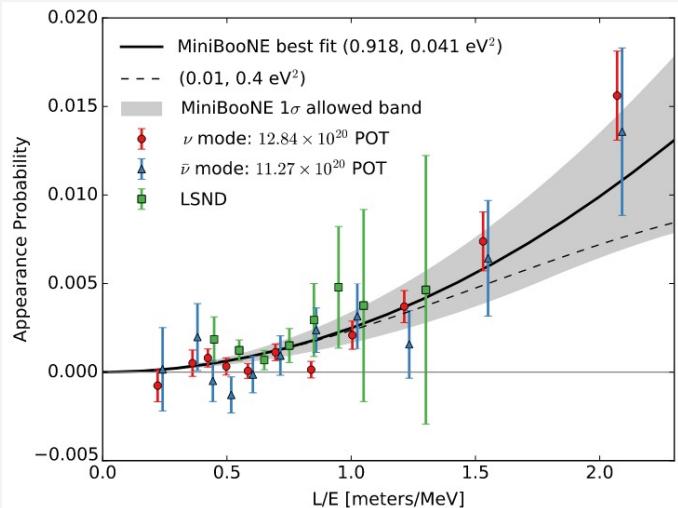
courtesy A. Göttel,
and 2104.02565

KM3NeT/ORCA is dedicated to solve the ordering too
6 strings operating, full operation expected in 2025,

R. Shanidze EPS HEP 2021

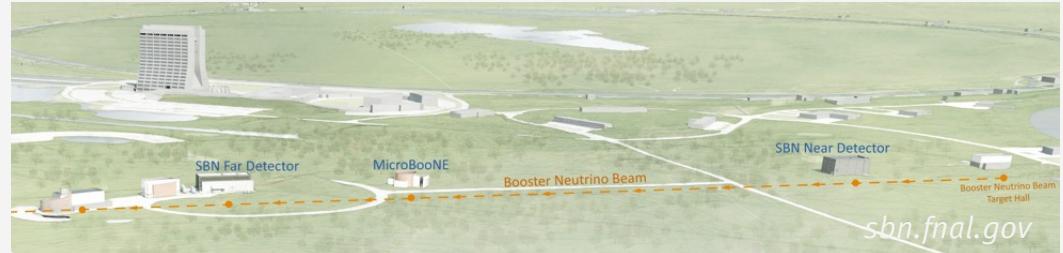
Short baseline program at Fermilab

Solve MiniBooNE $\nu_e, \bar{\nu}_e$ appearance signal
 “low energy excess of electromagnetic activity”



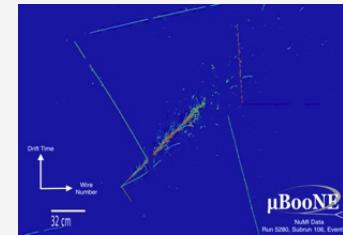
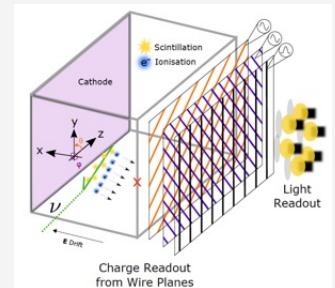
Continue developing LAr TPC technology

Short baseline program at Fermilab with LAr detectors

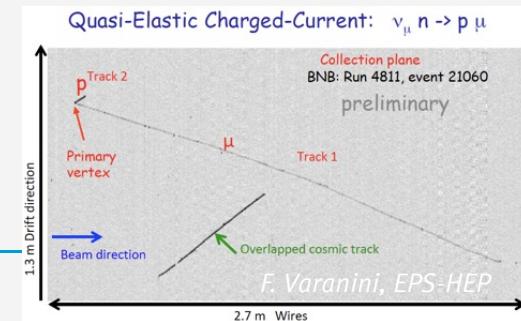


MicroBooNE measured flux-averaged inclusive $\nu_e, \bar{\nu}_e$ CC total cross section on Ar, excellent PID, e.g. γ/e^-

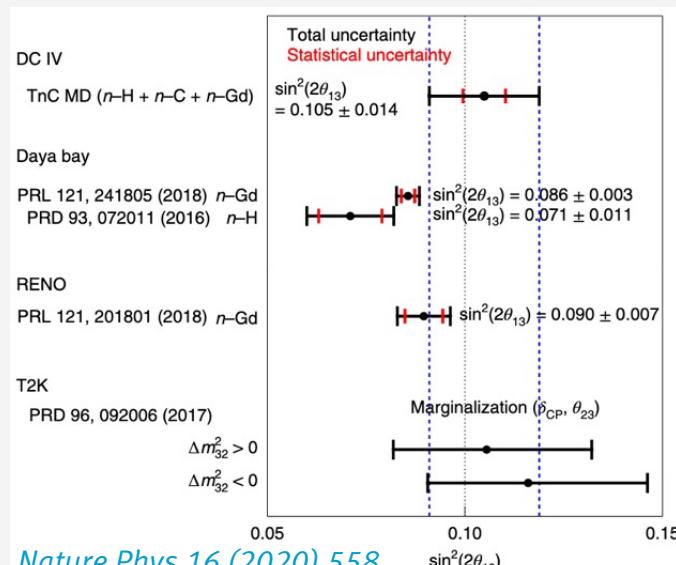
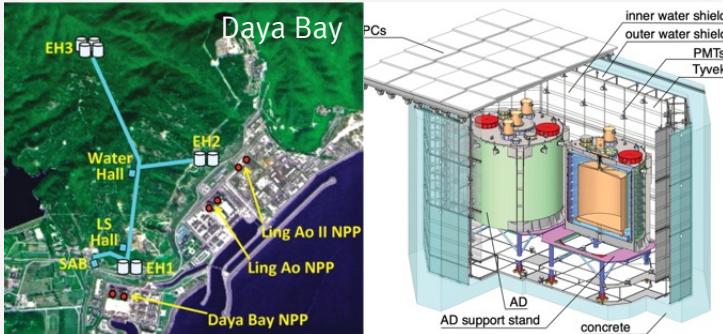
2101.04228



ICARUS has successfully taken first data



“Long baseline” reactor experiments: θ_{13} & more



Nature Phys 16 (2020) 558

$$\sin^2\theta_{13} = 0.0220 \pm 0.0007$$

(PDG2021 using Double Chooz, Reno, Daya Bay)

Checking LSND/MiniBooNE puzzle in 3+1 scenario
with Daya Bay and MINOS/MINOS+:

$$P_{\bar{\nu}_\mu \rightarrow \bar{\nu}_e}^{SBL} = 4|U_{e4}|^2|U_{\mu 4}|^2 \sin^2\left(\frac{\Delta m_{41}^2 L}{4E}\right)$$

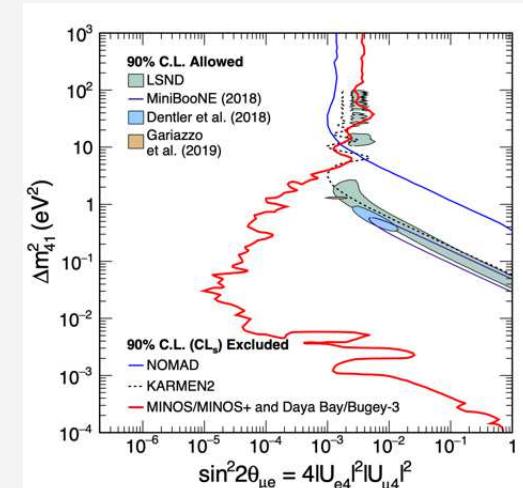
$$4|U_{e4}|^2|U_{\mu 4}|^2 = \sin^2 2\theta_{14} \sin^2 \theta_{24} \equiv \sin^2 2\theta_{\mu e}$$

Constrained by
 $\bar{\nu}_e$ disappearance
(Daya Bay and
Bugey-3)

Constrained by
 $\bar{\nu}_\mu$ disappearance
(MINOS & MINOS+)

→ exclusion of small
 Δm_{41} values

T. Schwetz at EPS HEP 2021:
"Sterile oscillation explanation of
LSND/MiniBooNE robustly disfavoured"



Phys Rev Lett 125 (2020) 07801

courtesy:
R. Mandujano

New: First MicroBooNE ν_e result

MicroBooNE at similar position in same ν beam as MiniBooNE

Data: February 2016 to July 2018, $7 \cdot 10^{20}$ POT in neutrino mode

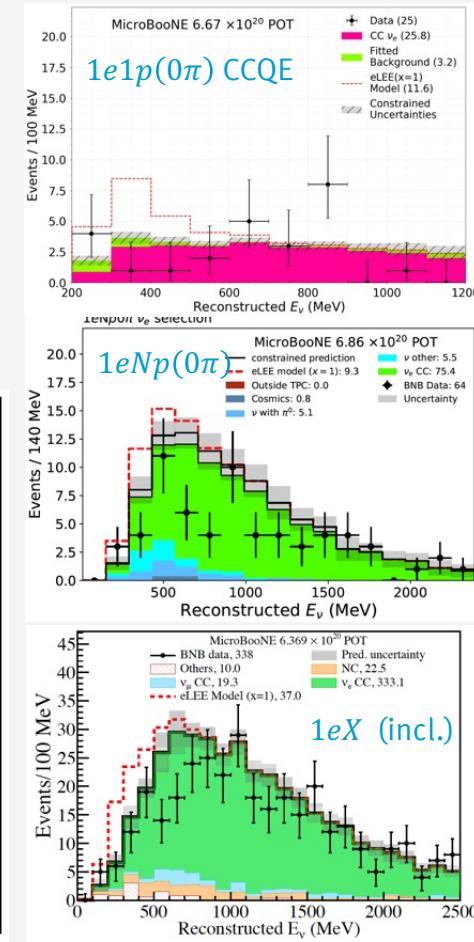
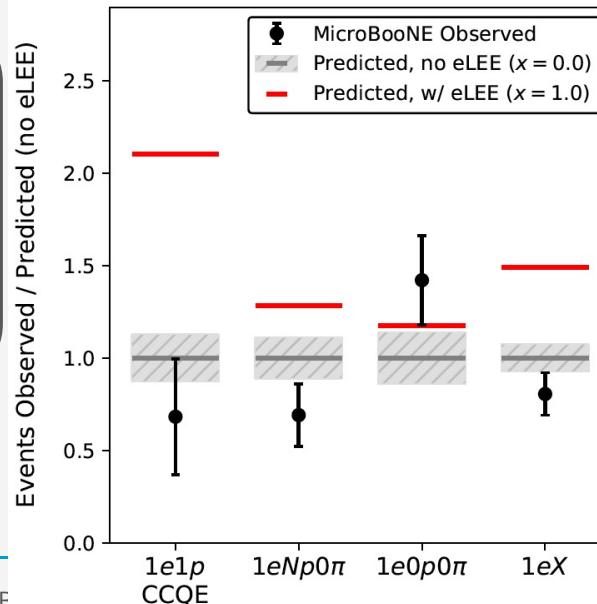
Different channels and different reconstruction methods

ν_e Final State	Signal Constraints	Reconstruction Approach
$1e1p(0\pi)$ CCQE	ν_μ CCQE	Deep-Learning [55]
$1eN(\geq 1)p0\pi, 1e0p0\pi$	ν_μ CC	Pandora [56]
$1eX$	ν_μ CC, ν_μ CC π^0 , ν_μ NC π^0	Wire-Cell [57]

No excess of ν_e with several inclusive and exclusive hadronic final states

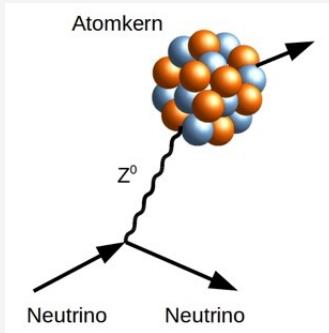
“MicroBooNE results are inconsistent with ν_e interpretation of MiniBooNE excess”

arXiv:2110.14054

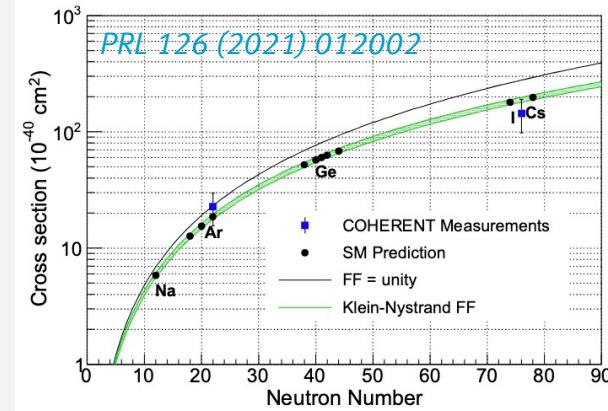
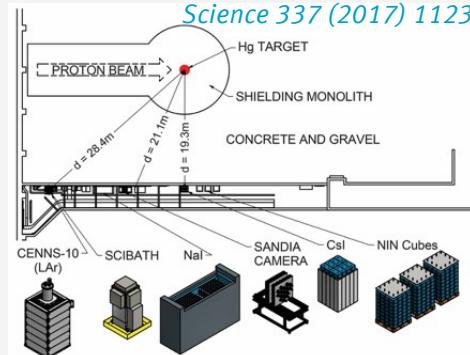


CEvNS: coherent elastic ν nucleon scattering

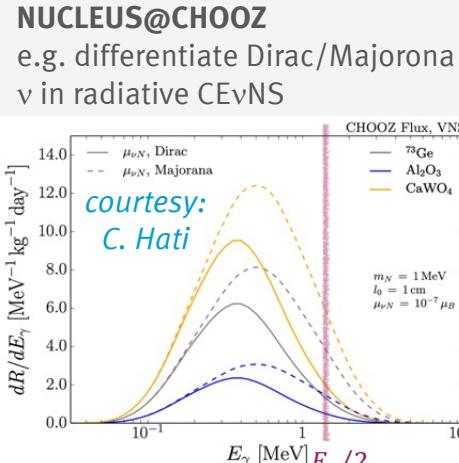
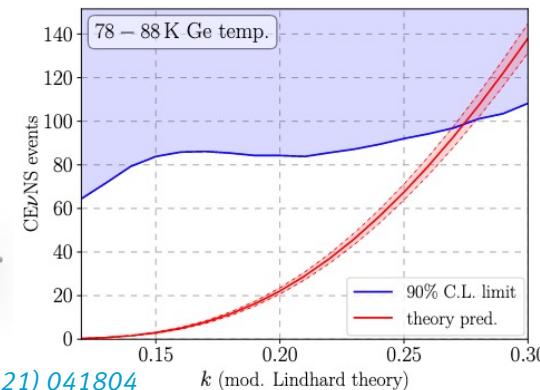
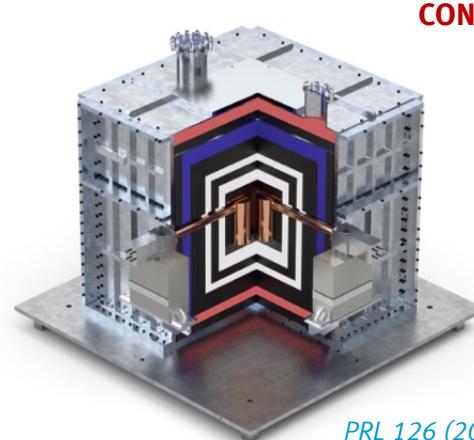
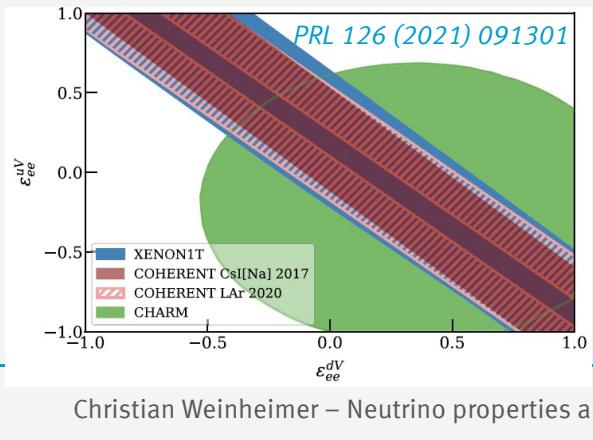
CEvNS observed on argon by COHERENT



www.mpi-hd.mpg.de



CEvNS open a new window for searching for BSM physics: NSI, Dirac vs Majorana



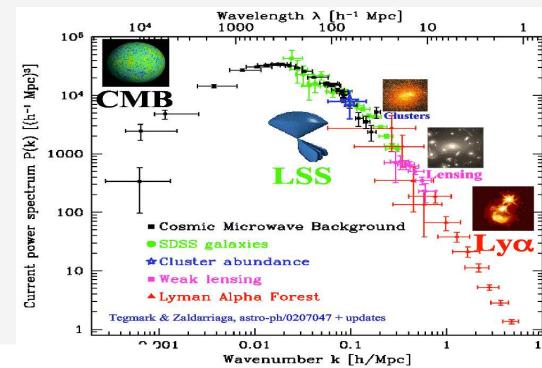
Three complementary ways to the absolute neutrino mass scale

1) Cosmology: $\sum_i m(\nu_i) = 3 \cdot \overline{m(\nu_i)}$

very sensitive, but model dependent, compares power at different scales

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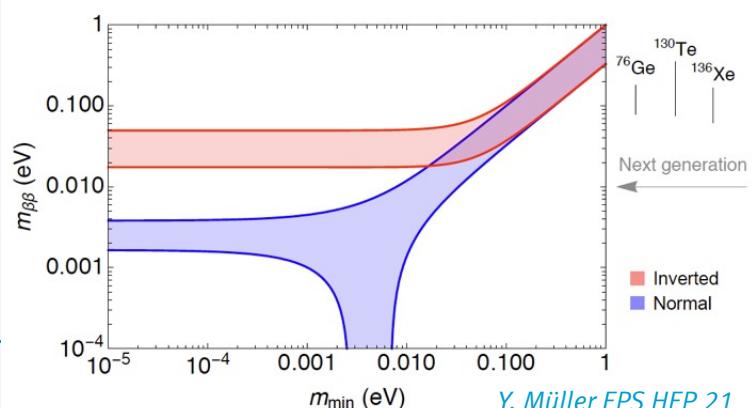
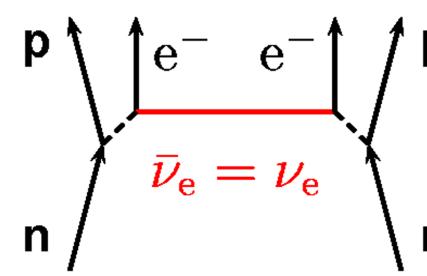
sensitive to Majorana neutrinos only, nuclear matrix elements

upper limits by EXO-200, KamLAND-Zen, GERDA, CUORE

disclaimer: $m_{\beta\beta}$ are valid only, if $0\nu\beta\beta$ works dominantly via ν exchange

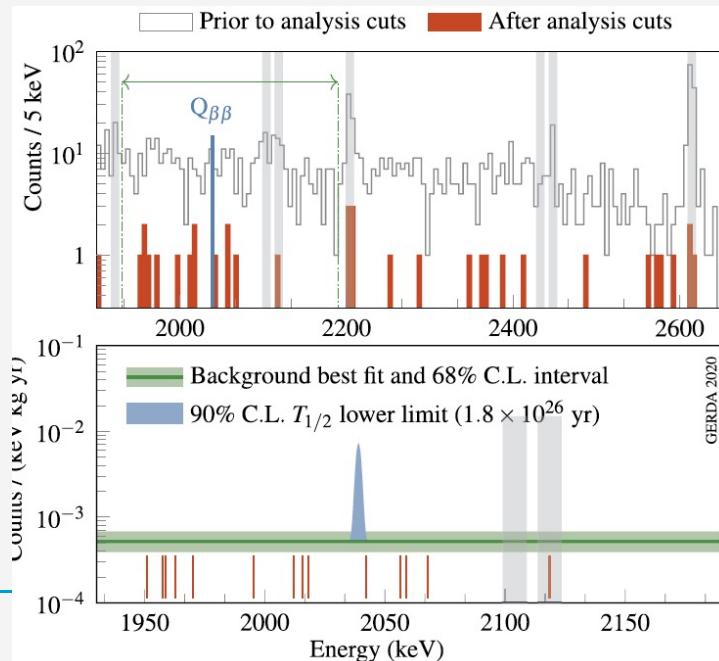
Discovery of $0\nu\beta\beta$ would proof lepton number violation !

$$T_{1/2}^{0\nu} \propto \begin{cases} a \cdot \varepsilon \cdot M \cdot t & \text{for background index } B = 0 \\ a \cdot \varepsilon \cdot \sqrt{\frac{M \cdot t}{B \cdot \Delta E}} & \text{for } B \neq 0 \end{cases}$$

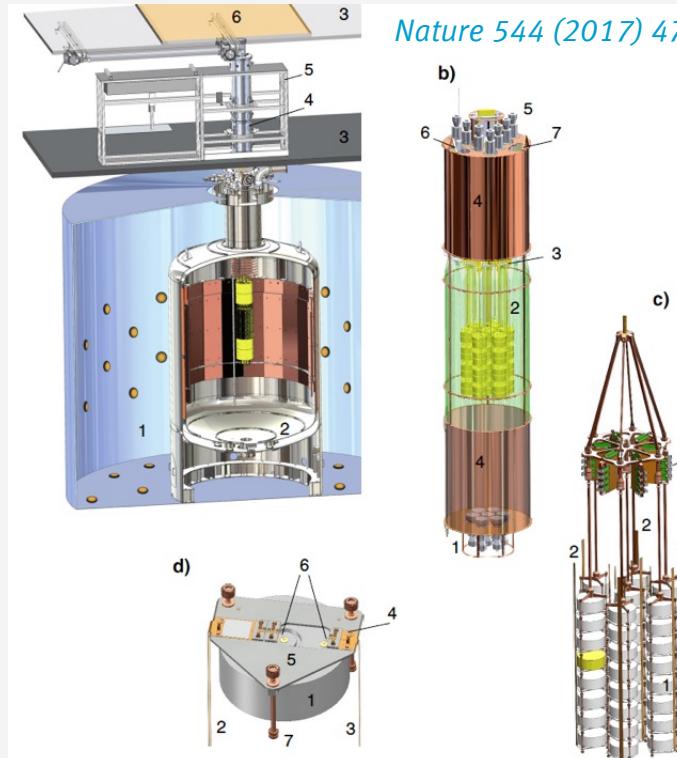
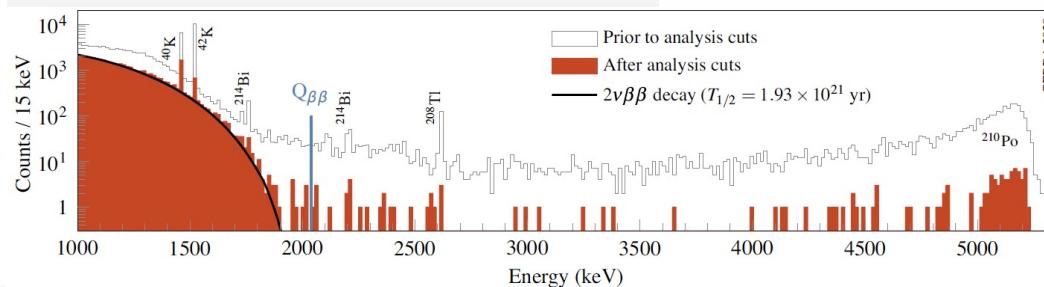


GERDA at LNGS – final result

enrichment of ^{76}Ge : $\approx 87\%$
 exposure: $127.2 \text{ kg}^*\text{yr}$
 energy resolution: $\approx 3 \text{ keV (FWHM)}$
 background index: $B = 5.2_{-1.3}^{+1.6} \cdot 10^{-4} \text{ counts}/(\text{keV kg yr})$
→ lower limit: $T_{1/2}^{0\nu} > 1.8 \cdot 10^{26} \text{ yr}$ (90% C.L.)
→ upper mass limit: $m_{\beta\beta} < 79 - 180 \text{ meV}$ (without g_A quenching)



PRL 125 (2020) 252502



^{76}Ge : LEGEND-200 at LNGS and LEGEND-1000

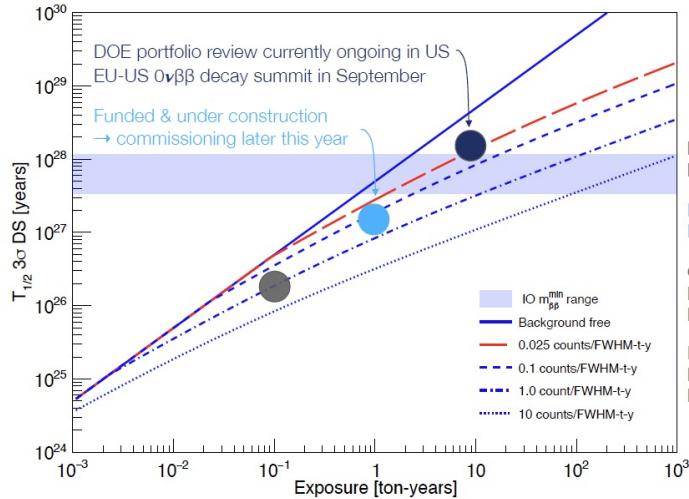
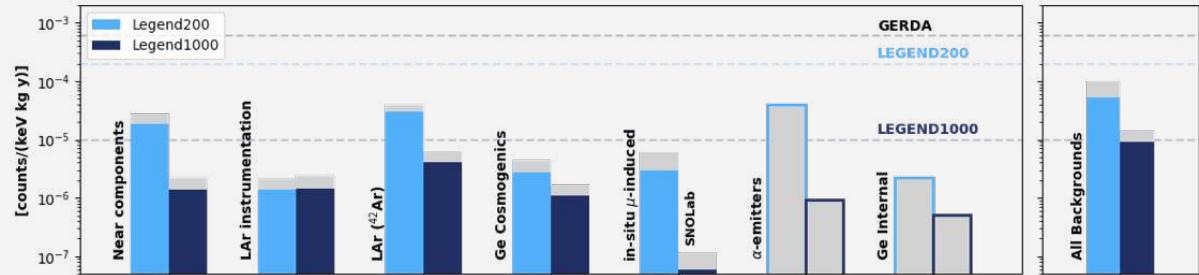
based on GERDA+Majorana technologies

enriched ^{76}Ge , PPC/BEGe/ICPC detectors (\rightarrow event topology)

LEGEND-200: start data taking (140-150 kg) still in 2021

upgrade by another 50 kg in 2022

LEGEND -1000: future plan, PCDR 2107.11462



Detector mass: 1000 kg $T_{1/2}^{0\nu} > 10^{28}$ yr
Exposure: 10 ton-yr

Detector mass: 200 kg $T_{1/2}^{0\nu} > 10^{27}$ yr
Exposure: 1 ton-yr

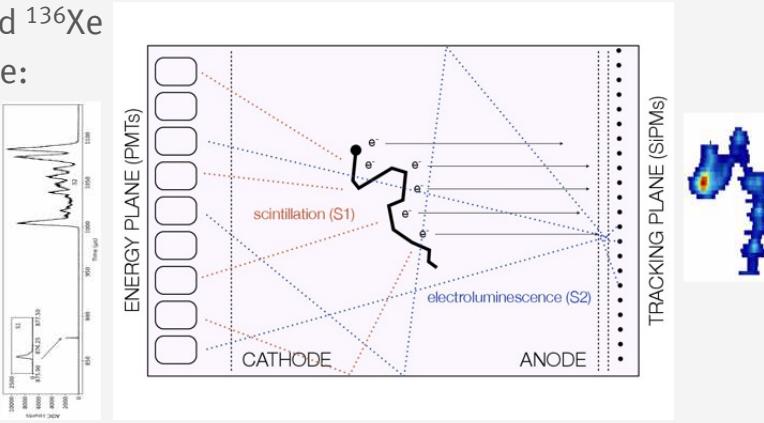
GERDA:
Detector mass: 44 kg $T_{1/2}^{0\nu} > 1.8 \cdot 10^{26}$ yr
Exposure: 127 kg-yr

MAJORANA DEMONSTRATOR:
Detector mass: 30 kg $T_{1/2}^{0\nu} > 2.7 \cdot 10^{25}$ yr
Exposure: 26 kg-yr

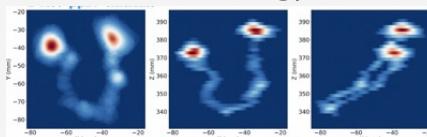
Future of xenon-based $0\nu\beta\beta$ experiments

NEXT: GXe TPC

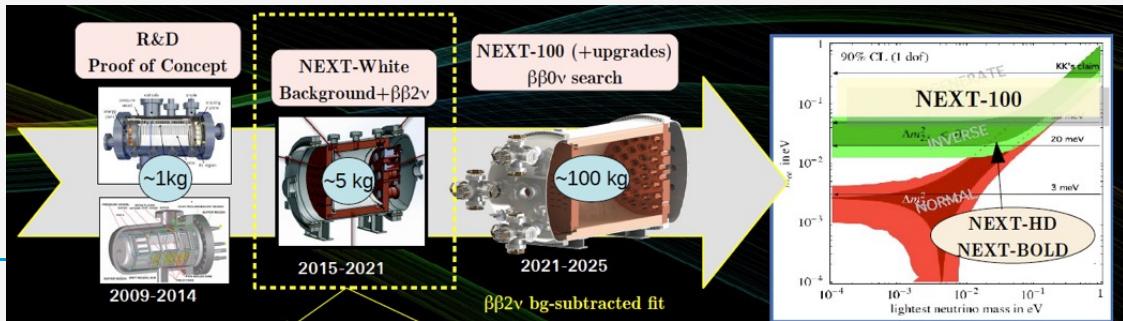
enriched ^{136}Xe
principle:



advantage of GXe TPC: 2-times better energy resolution
topology of $\beta\beta$ event



P. Novella, EPS HEP 2021



nEXO: LXe TPC

enriched ^{136}Xe :

5 t

energy resolution:

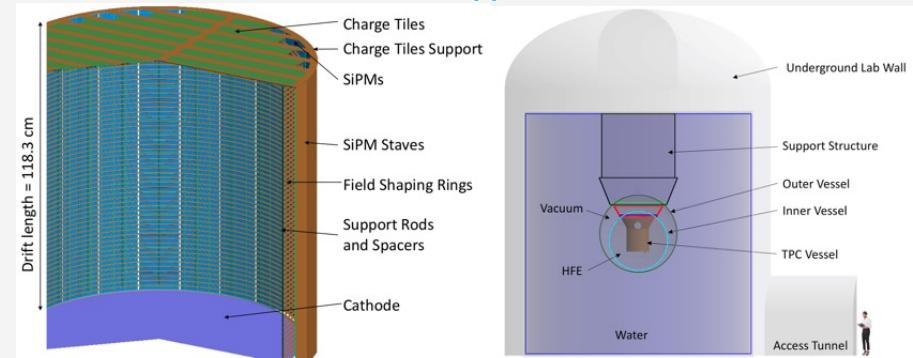
$\approx 46 \text{ keV (FWHM)}$

background index:

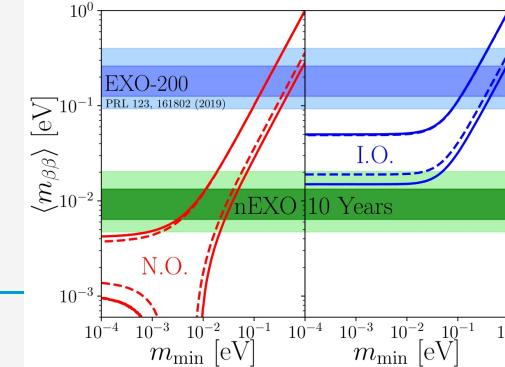
$B = 7 \cdot 10^{-5} \text{ counts/(FWHM kg yr)}$

→ expected sensitivity (10 yr): $T_{1/2}^{0\nu} > 1.35 \cdot 10^{28} \text{ yr (90\% C.L.)}$

→ expected sensitivity (10 yr): $m_{\beta\beta} < 5 - 20 \text{ meV}$



2106.16243



The dark matter experiments DARWIN will have a sensitivity of $2.4 \cdot 10^{27} \text{ yr}$, EPJC 80 (2020) 808
J. Masbou, EPS HEP 2021

CUPID in CUORE cryostat at LNGS

enrichment of ^{100}Mo : 95%

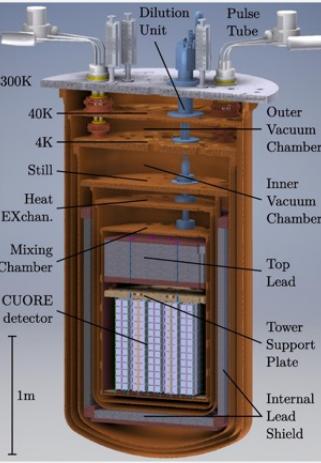
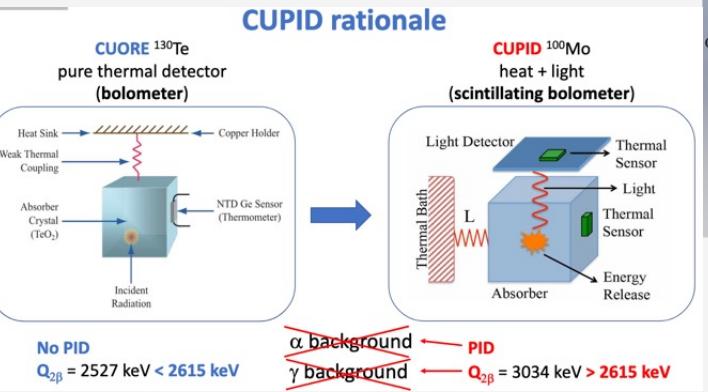
mass (^{100}Mo) in $\text{Li}_2^{100}\text{MoO}_4$: 240 kg

energy resolution: 5 – 7 keV (FWHM)

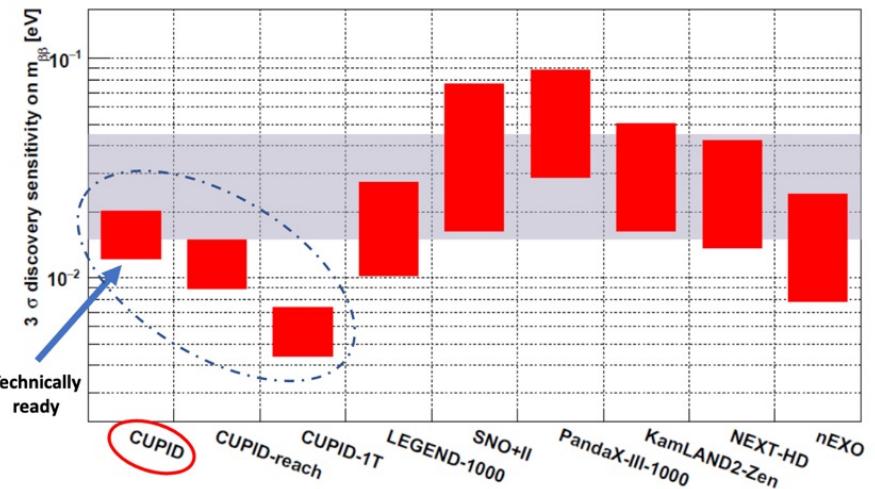
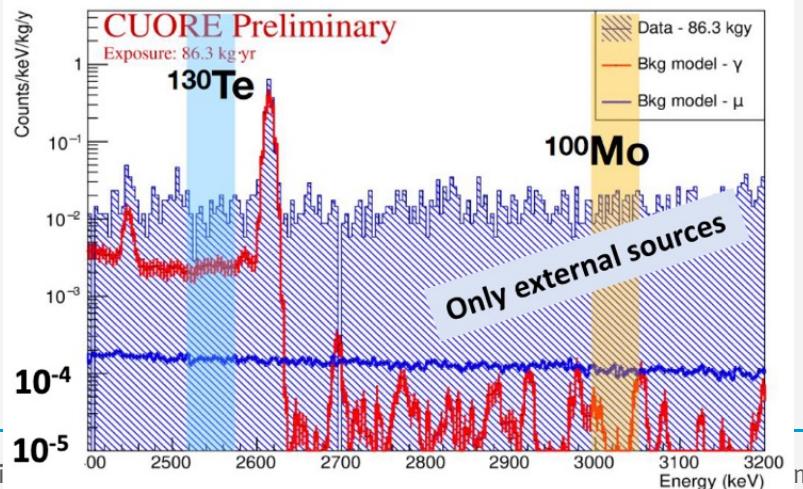
background index: $B = 1 \cdot 10^{-4}$ counts/(keV kg yr)

→ exp. sensitivity (10yr): $T_{1/2}^{0\nu} > 1.4 \cdot 10^{27}$ yr (90% C.L.)

→ exp. sensitivity (10yr): $m_{\beta\beta} < 10 - 17$ meV



A. Giuliani, Neutrino Telescopes 2021



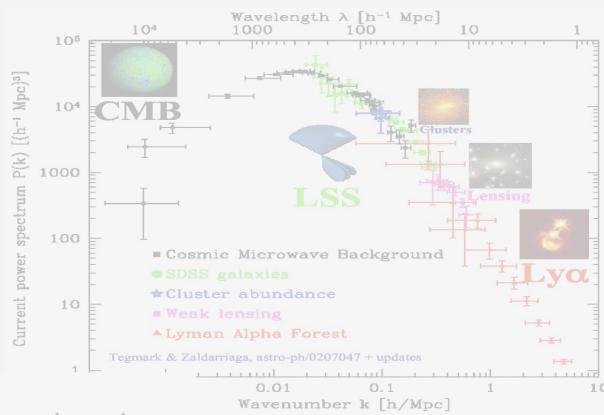
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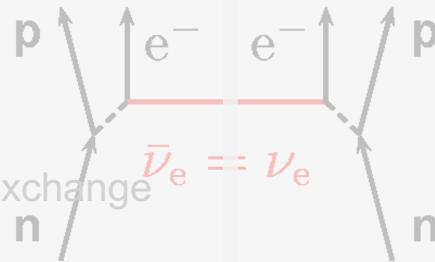
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sensitive to Majorana neutrinos only, nuclear matrix elements

upper limits by EXO-200, KamLAND-Zen, GERDA, CUORE

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discovery of $0\nu\beta\beta$ would mean lepton number violation !



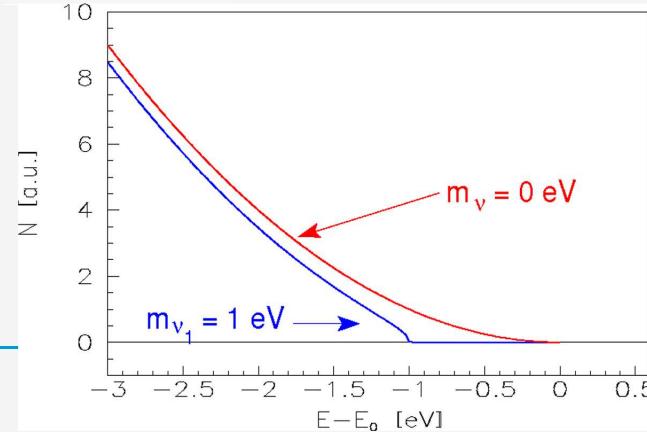
3) Direct neutrino mass determination: $m^2(\nu_e) := m_\beta^2 := \sum_i |U_{ei}|^2 \cdot m^2(\nu_i)$

no further assumptions needed, use $E^2 = p^2c^2 + m^2c^4 \rightarrow m^2(\nu)$

Time-of-flight measurements (ν from supernova)

Kinematics of weak decays / beta decays, e.g. tritium, ^{163}Ho

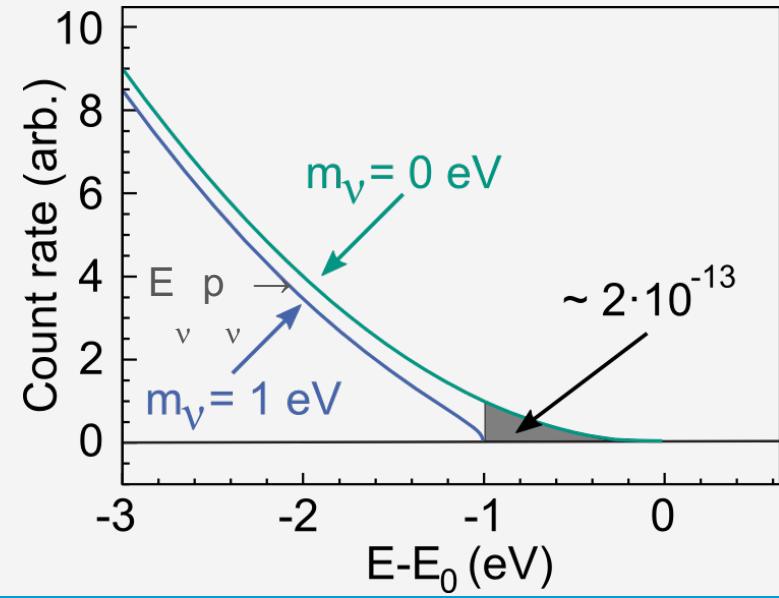
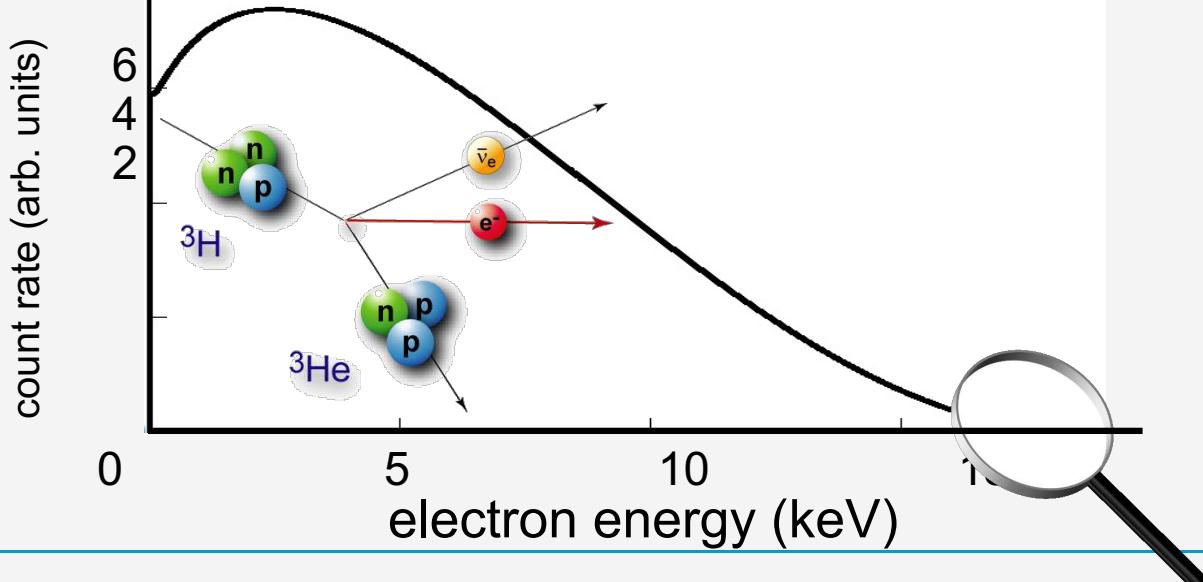
measure charged decay prod., E-, p-conservation



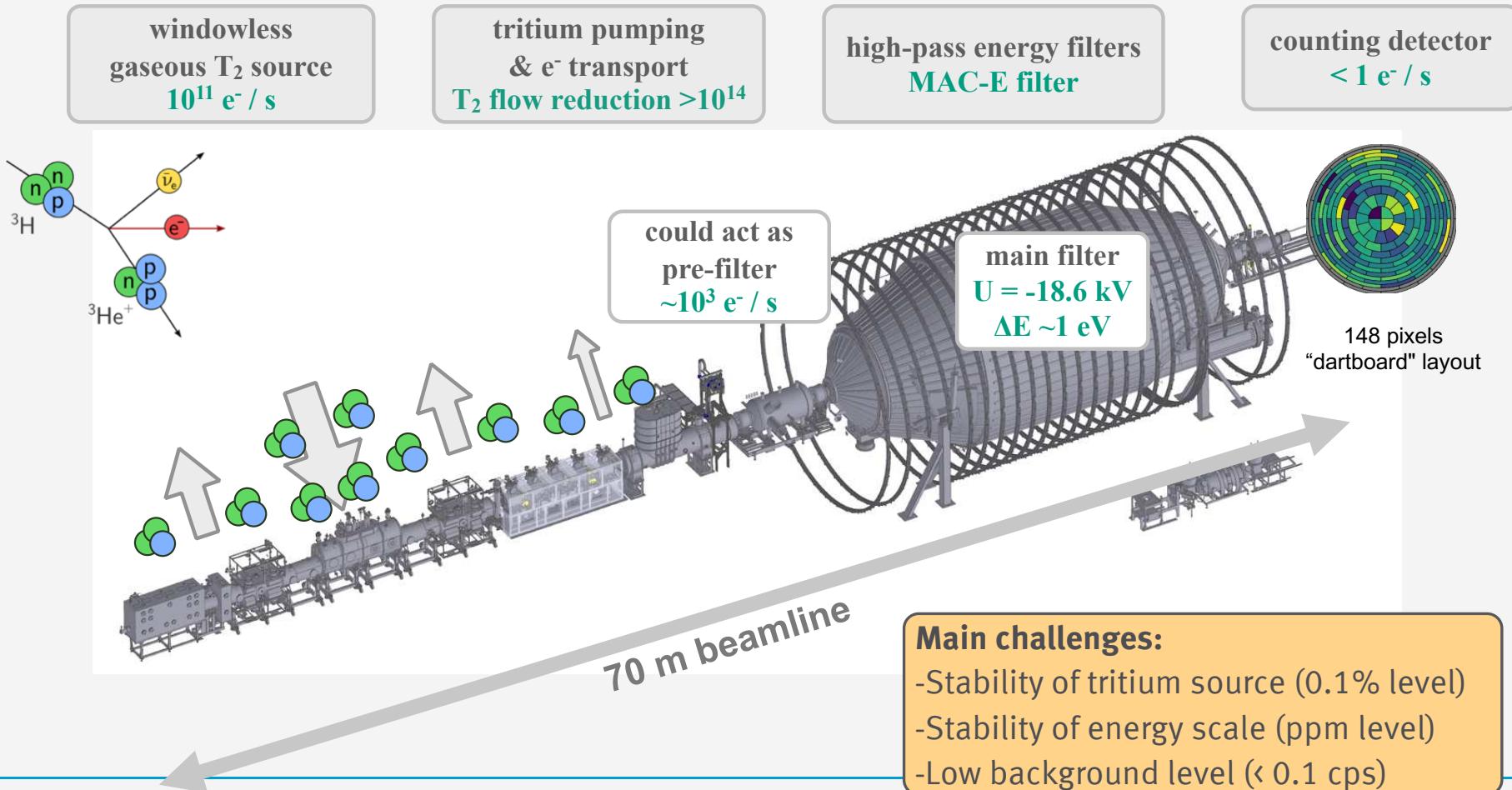
Direct determination of "m(ν_e)" from β -decay (EC)

$$\text{β: } \frac{dN}{dE} = K \cdot F(E, Z) \cdot p \cdot E_{tot} \cdot (E_0 - Ee) \cdot \underbrace{\sum_i |U_{ei}|^2}_{\text{essentially phase space: } p_e} \cdot \underbrace{\sqrt{(E_0 - Ee)^2 - m^2(\nu_i)}}_{E_\nu} \cdot \underbrace{p_\nu}_{p_\nu}$$

with "electron neutrino mass": " $m^2(\nu_e)$ " $\coloneqq \sum_i |U_{ei}|^2 \cdot m^2(\nu_i)$, complementary to $0\nu\beta\beta$ & cosmology
(modified by electronic final states, recoil corrections, radiative corrections)



The KArlsruhe TRItium Neutrino experiment KATRIN



The KArlsruhe TRItium Neutrino experiment KATRIN



The international KATRIN Collaboration: 150 people from 20 (6) institutions (countries)



Funded by:



Bundesministerium
für Bildung
und Forschung



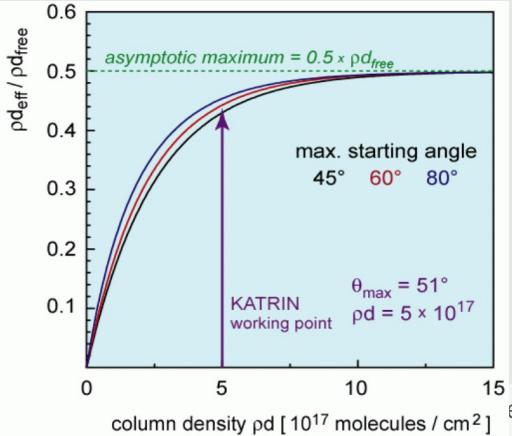
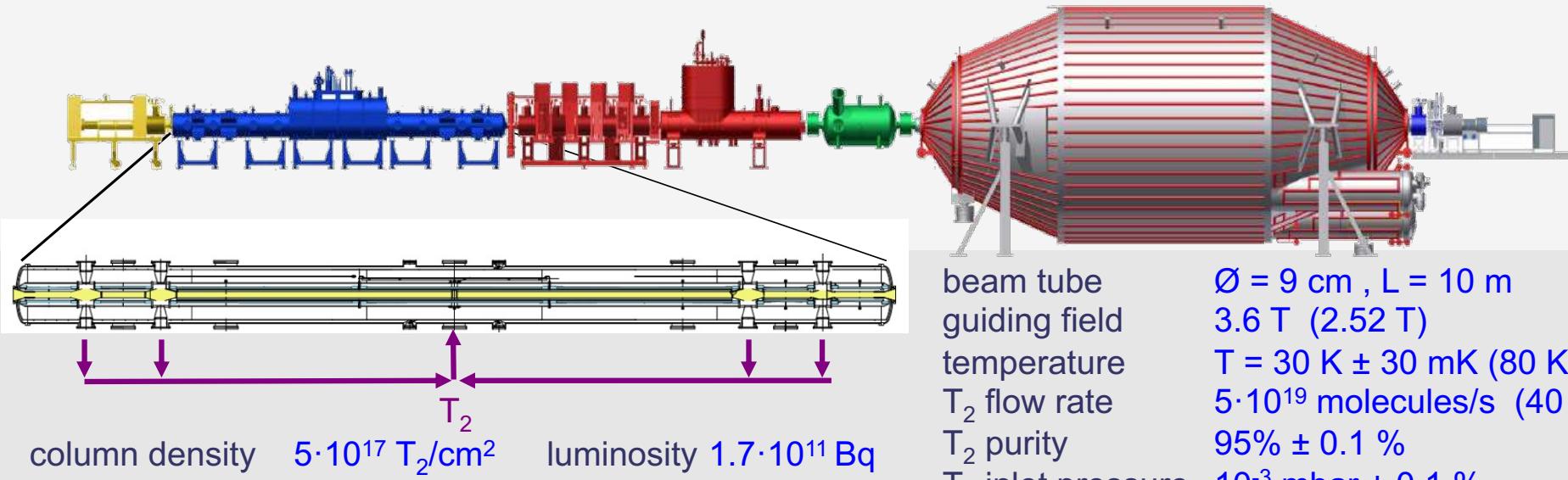
Czech
Republic:



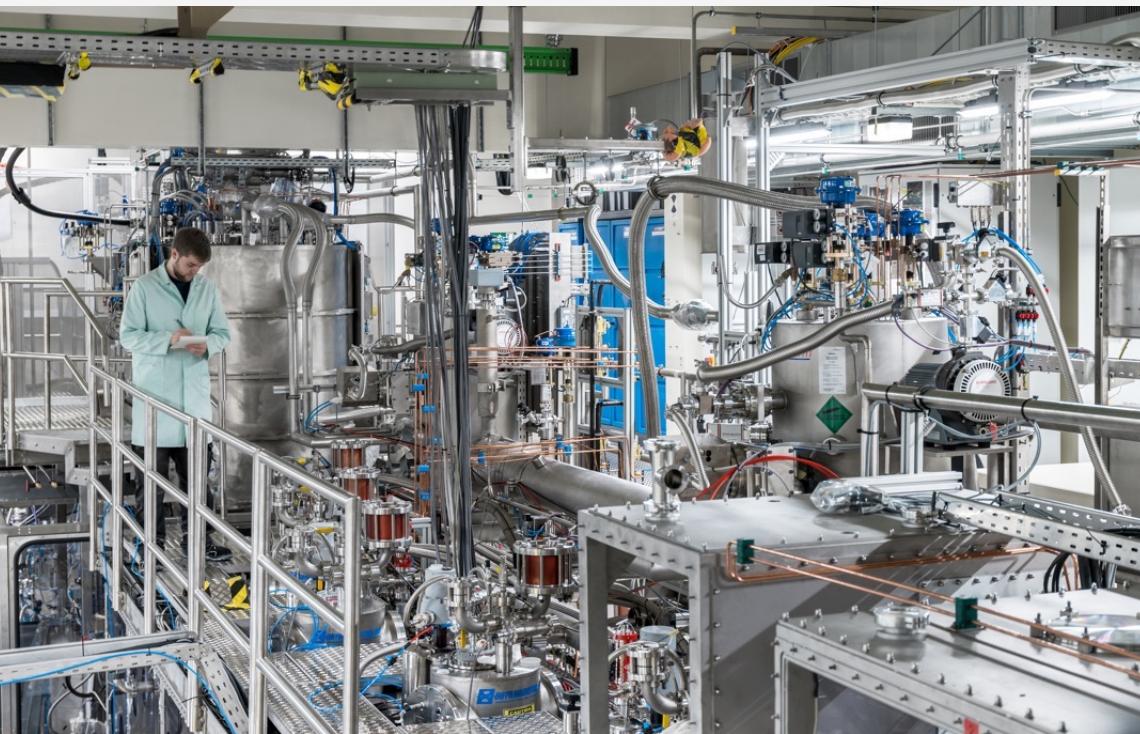
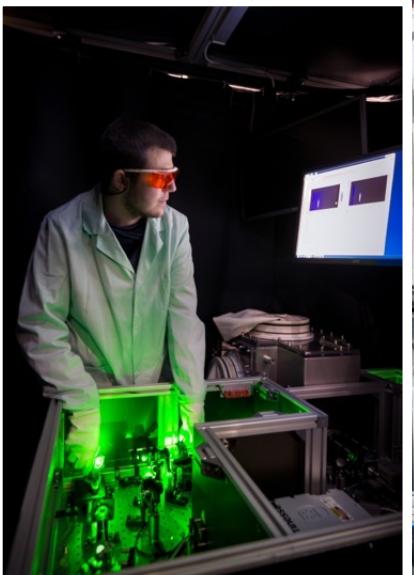
MINISTRY OF EDUCATION
YOUTH AND SPORTS



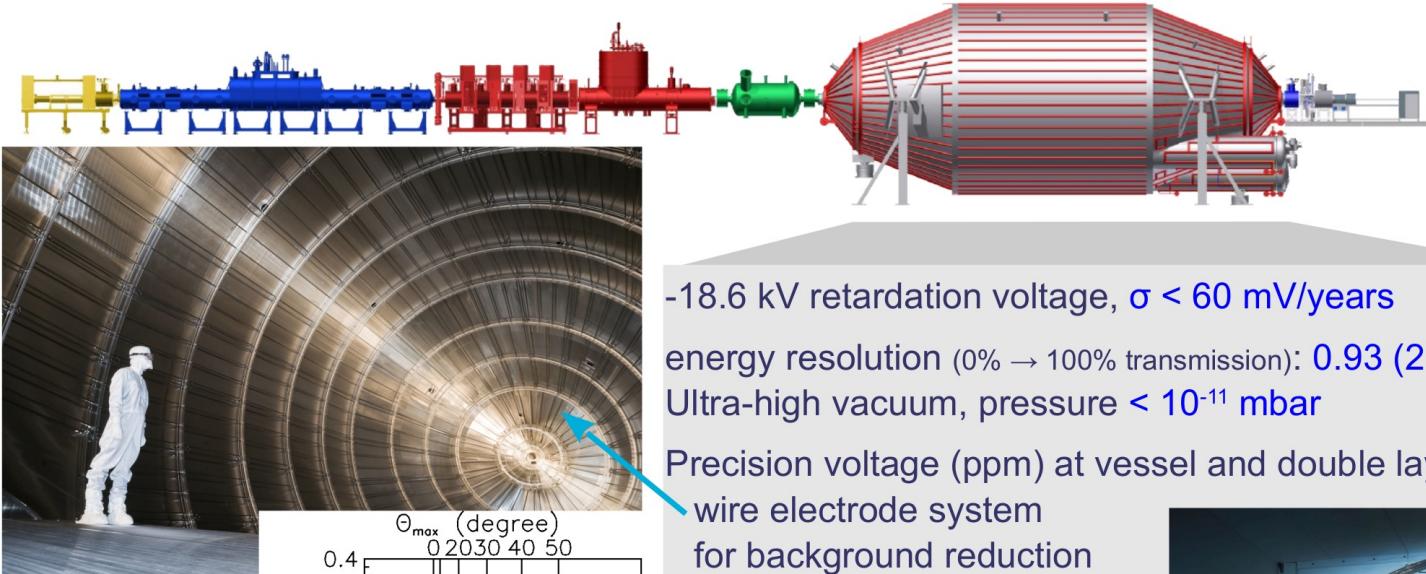
The Windowless Gaseous Molecular Tritium Source



Photos: source & transport section

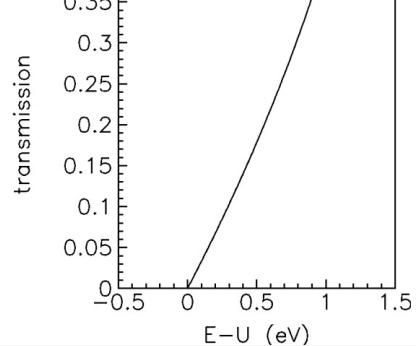


The main spectrometer: an integrating high resolution MAC-E Filter



→ integral transmission function:

$$\Delta E = E \cdot B_{\min} / B_{\max} = 0.93 \text{ eV} \quad (2.7 \text{ eV})$$



-18.6 kV retardation voltage, $\sigma < 60 \text{ mV/years}$
 energy resolution (0% → 100% transmission): 0.93 (2.7) eV
 Ultra-high vacuum, pressure $< 10^{-11} \text{ mbar}$

Precision voltage (ppm) at vessel and double layer
 wire electrode system
 for background reduction
 and field shaping

Air coils for earth magnetic
 field compensation



The KATRIN electron detector FPD

Focal plane detection system

segmented Si PIN diode:

90 mm Ø, 148 pixels, 50 nm dead layer

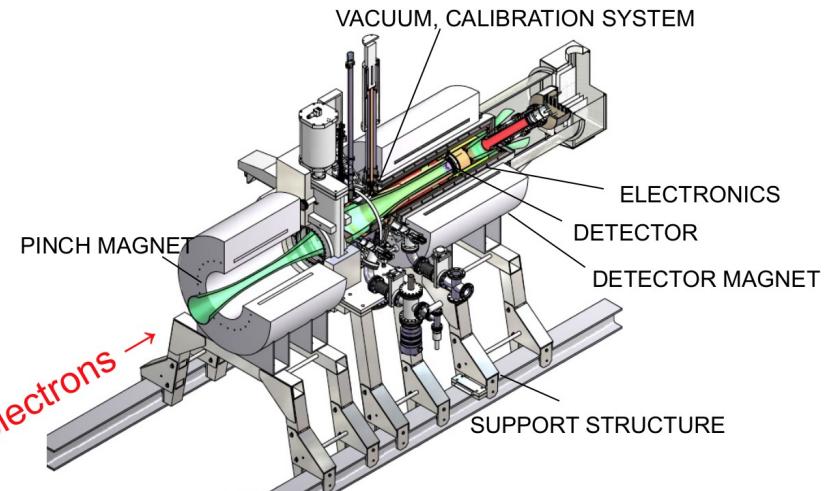
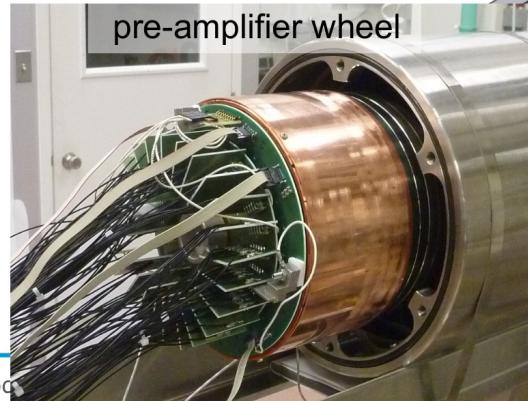
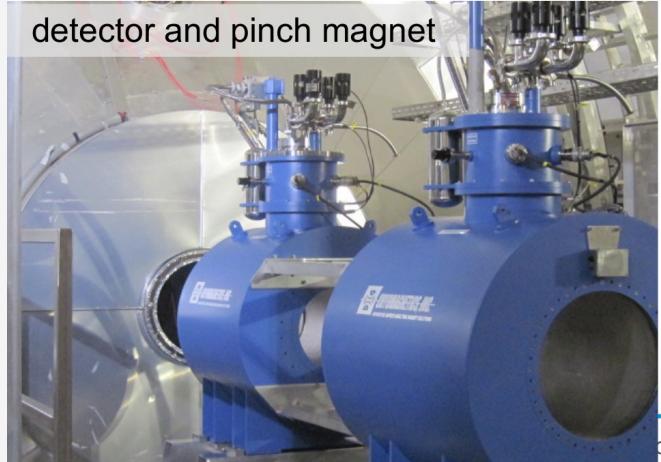
energy resolution ≈ 1 keV

pinch and detector magnets up to 6 T

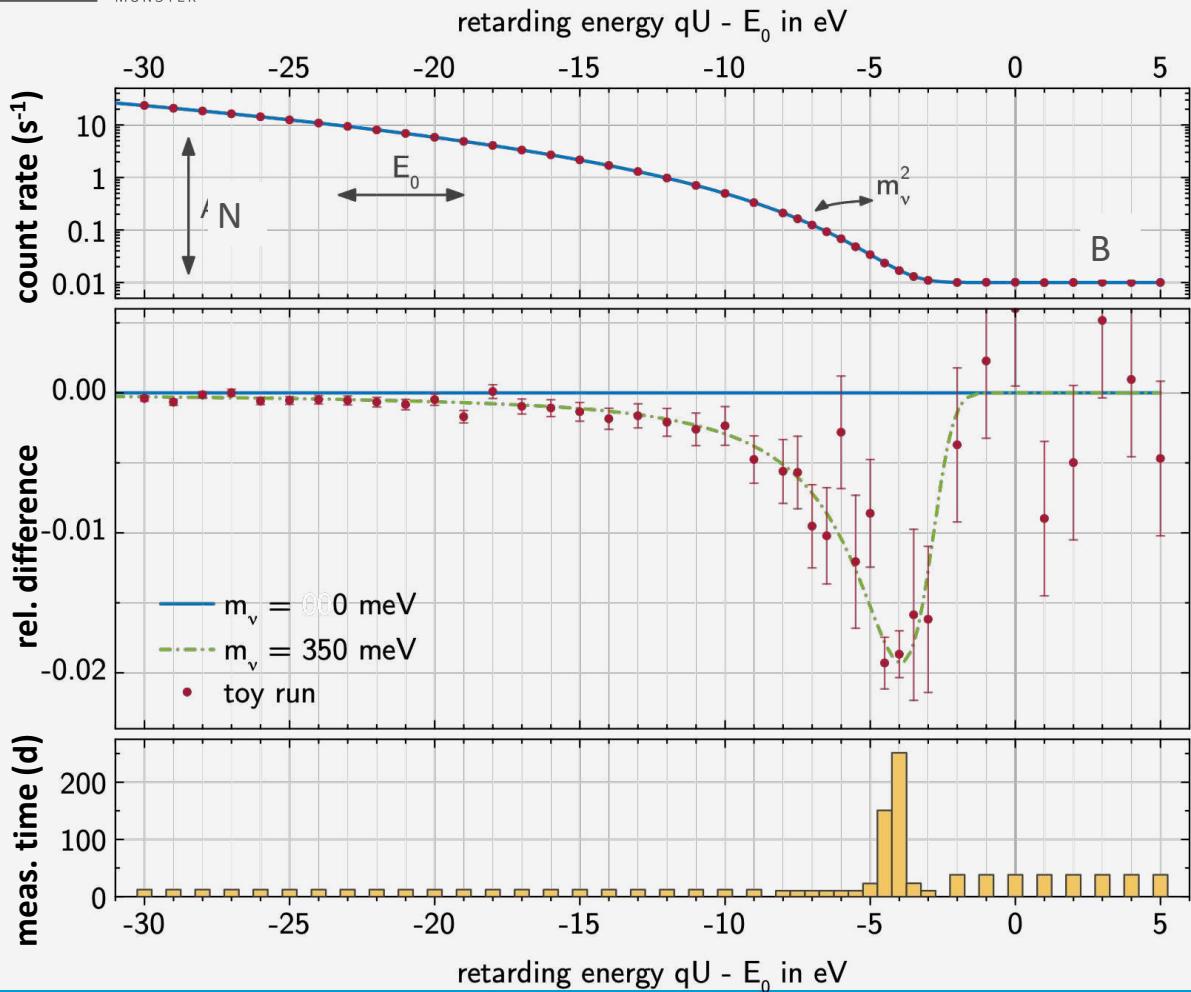
post acceleration (10kV)

active veto shield

detector and pinch magnet



The measurement principle



Direct **shape** measurement
of **integrated β spectrum**
Four fit parameters:

spectrum
norm. **N**

spectrum
endpoint **E_0**

background
rate **B**

squared
mass **m_ν^2**

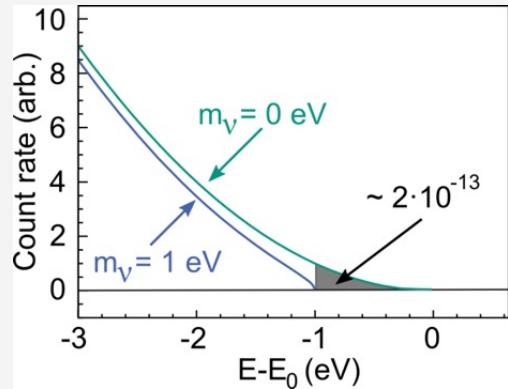


$\sim 10^{-8}$ of all β -decays in scan region
 ~ 40 eV below endpoint

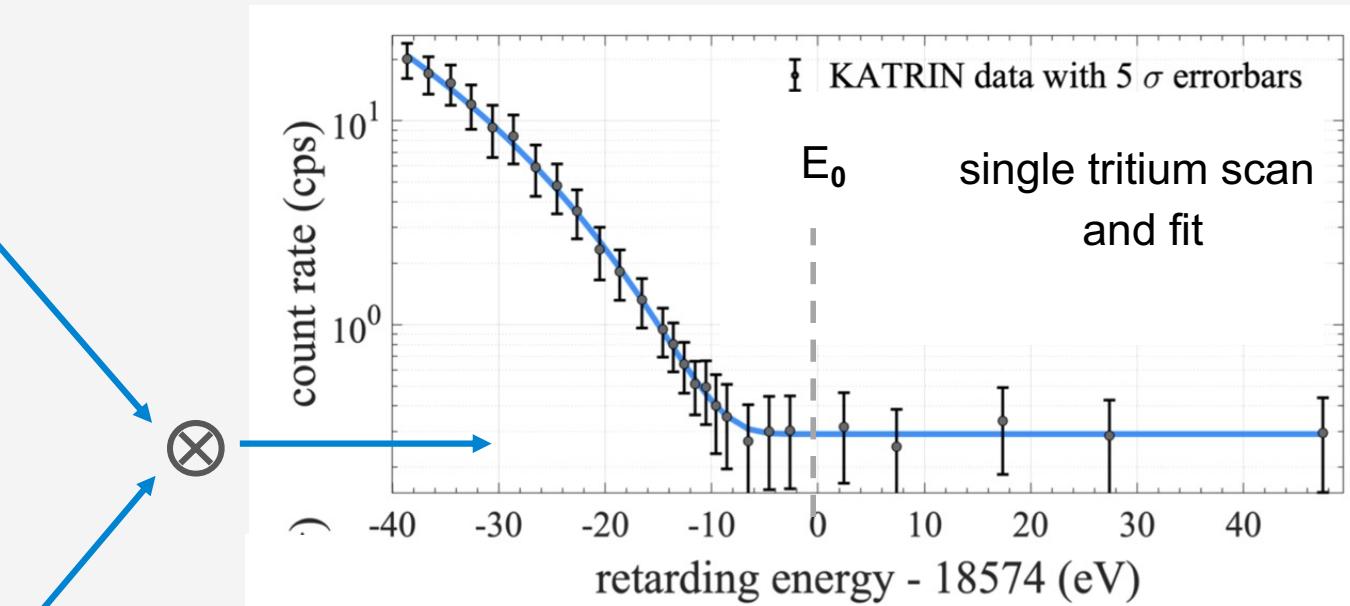
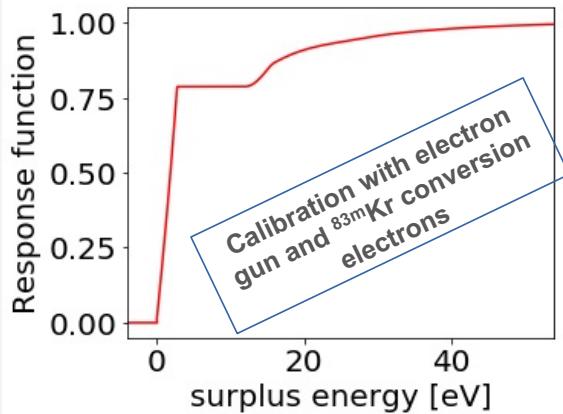
EPJ C 79 (2019) 204

Model of the experimental spectrum

- Beta spectrum: $R_\beta(E, m^2(\nu_e))$



- Experimental response: $f(E - qU)$



$$R(qU) = A_s \cdot N_T \int_{qU}^{E_0} R_\beta(E, m^2(\nu_e)) \cdot f(E - qU) dE + R_{bg}$$

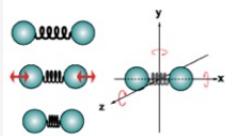
PRL 123 (2019) 221802, EPJ C 79 (2019) 204
+ detailed analysis PRD 104 (2021) 012005
+ energy loss measurement EPJ C 81 (2021) 579

Systematic effects and uncertainties



Molecular final states

- quantum-chemical
- computations



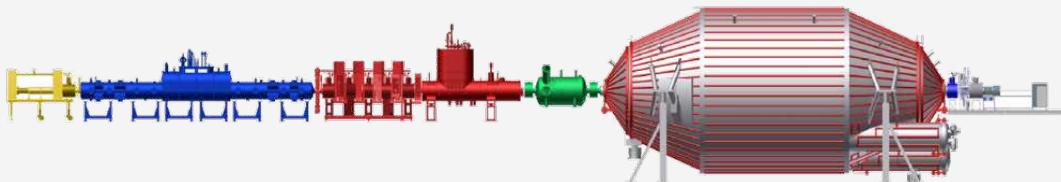
Source electric potential

- plasma properties
- surface conditions



Magnetic fields

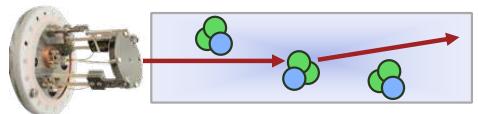
- source
- spectrometer
- detector



Detection efficiency



Energy loss by scattering

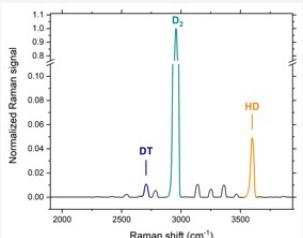


EPJ C 79 (2019) 204
EPJ C 81 (2021) 579

Activity fluctuations

- column density
- tritium (T_2 , DT, HT)
- concentration

Sensors 20 (2020) 4827



Background

- dependence on retarding potential
- time structure due to trapped electrons

arXiv:2011.05107

Three complementary strategies to include systematics in the fit:

(a) covariance matrix, (b) Monte-Carlo propagation, (c) pull-term method

see *PRL* 123 (2019) 221802 + detailed analysis *PRD* 104 (2021) 012005

First neutrino mass result

1st science run of KATRIN in spring 2019

- ν-mass: best fit result

$$m^2(\nu_e) = -1.0^{+0.9}_{-1.1} \text{ eV}^2$$

- ν-mass: new upper limit

$$m(\nu_e) < 1.1 \text{ eV (90% C.L. Lokhov – Tkachov)}$$

$m(\nu_e) < 0.8 \text{ eV}$ (90% C.L. Feldman-Cousins)

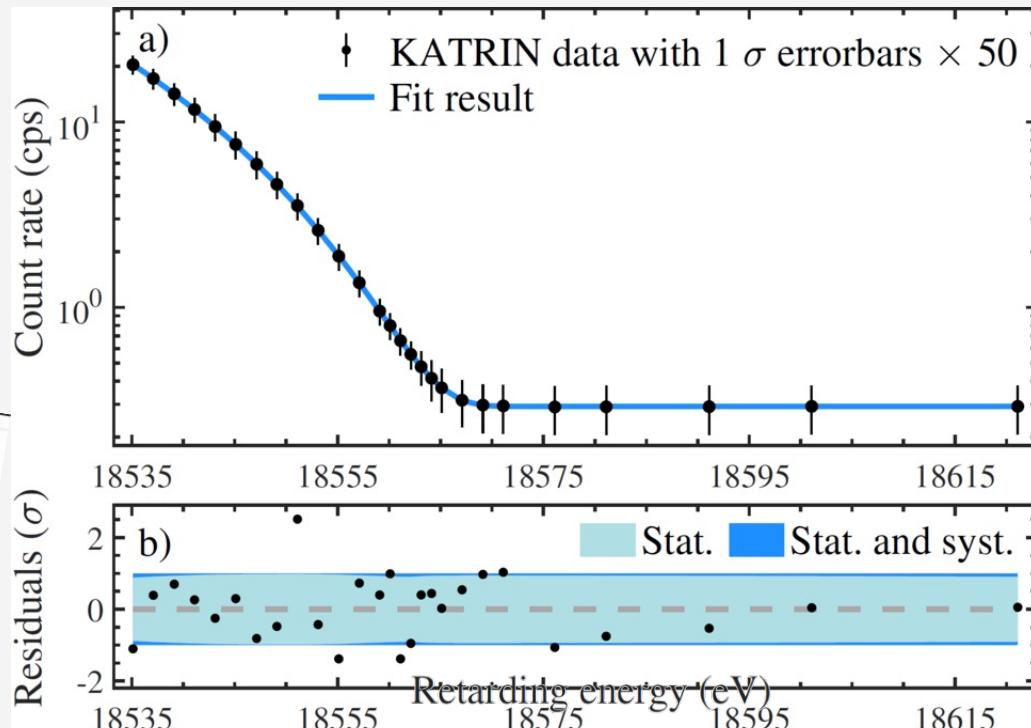
$m(\nu_e) < 0.9 \text{ eV}$ (90% C.L. Bayesian, flat prior $m^2 > 0$)

Excellent data quality

Compared to previous experiments:

- Improvement of statistics (x2)
- Reduction of systematics (÷6)

Only 9 days out of 1000 live days!

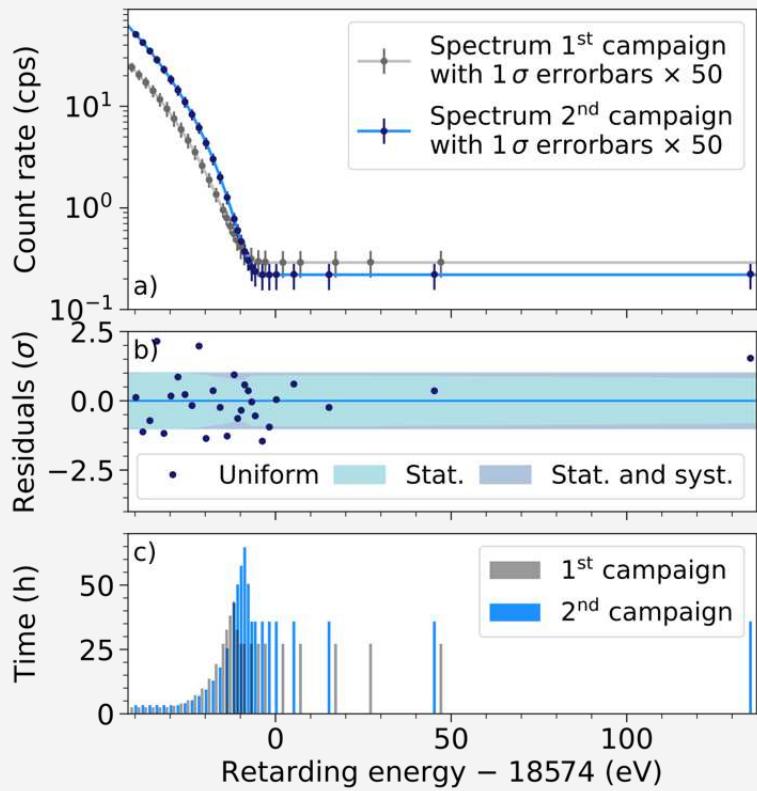


PRL 123 (2019) 221802 + detailed analysis PRD 104 (2021) 012005

Improvements of 2nd campaign compared to 1st one

	1 st campaign PRL 123 (2019) 221802	2 nd campaign This talk, arXiv:2105.08533
Campaign date	April-May 2019	Sept-Nov 2019
Total scan time	522 h (274 scans)	744 h (361 scans)
Source activity	25 GBq	98 GBq
Background	290 mcps	220 mcps
Tritium purity	97.6%	98.7%
Electrons in R ₀	2 Mio	4.3 Mio

2nd KATRIN neutrino mass campaign



PRL 123 (2019) 221802
arXiv:2105.08533 (subm. to Nature Physics)

$$m^2(\nu) = (0.26^{+0.34}_{-0.34}) \text{ eV}^2$$

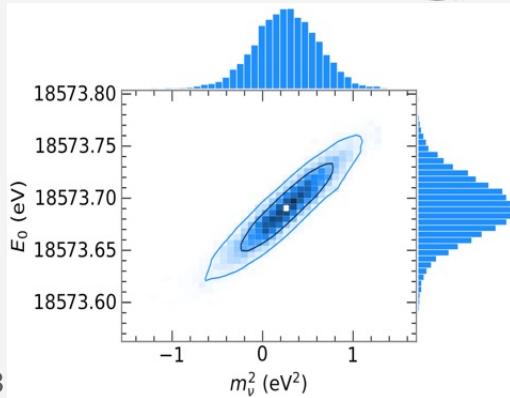
→ compatible with zero

$$E_0 = 18573.69 \pm 0.03 \text{ eV}$$

→ Q-value : $18575.2 \pm 0.5 \text{ eV}$

good agreement with Penning trap exp.:

$Q = 18575.72 \pm 0.07 \text{ eV}$, PRL 114 (2015) 013003



Frequentist limit: $m_\nu < 0.9 \text{ eV}$ (90% CL)

Same for Feldman & Cousins and Lokhov & Tkachov
less than sensitivity, due to positive fit result

2105.08533

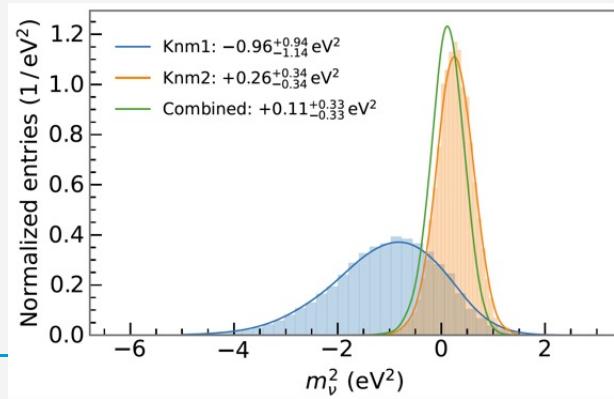
Bayesian: $m_\nu < 0.85 \text{ eV}$ (90% CI)

Lokhov & Tkachov, Phys. Part. Nucl. 46 (2015) 347
Feldman & Cousins, Phys. Rev. D57 (1998) 3873

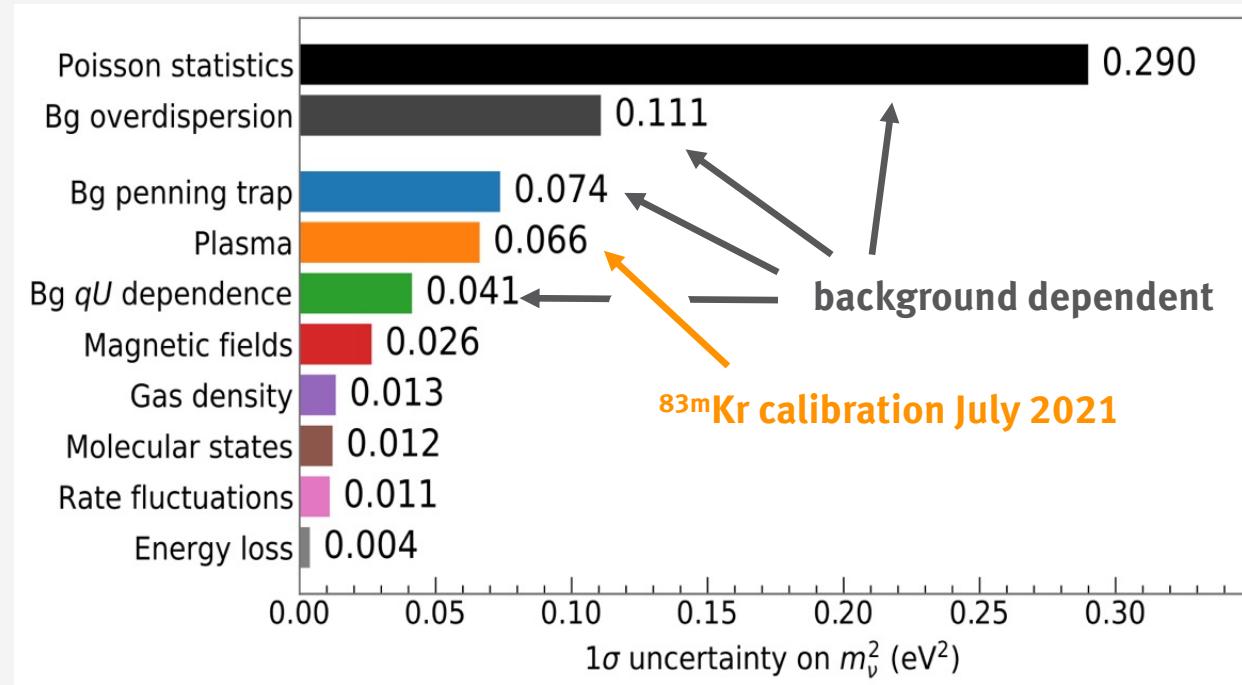
Combine 1st & 2nd campaign:

Freq. im.: $m_\nu < 0.8 \text{ eV}$ (90% CL)

Bayes: $m_\nu < 0.7 \text{ eV}$ (90% CL)

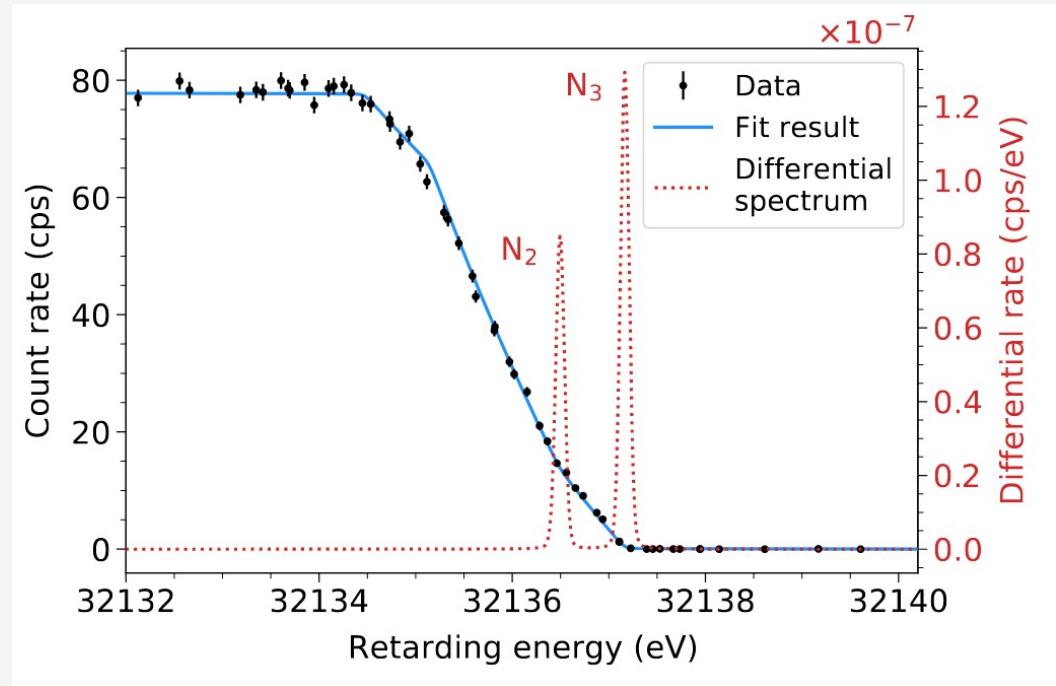
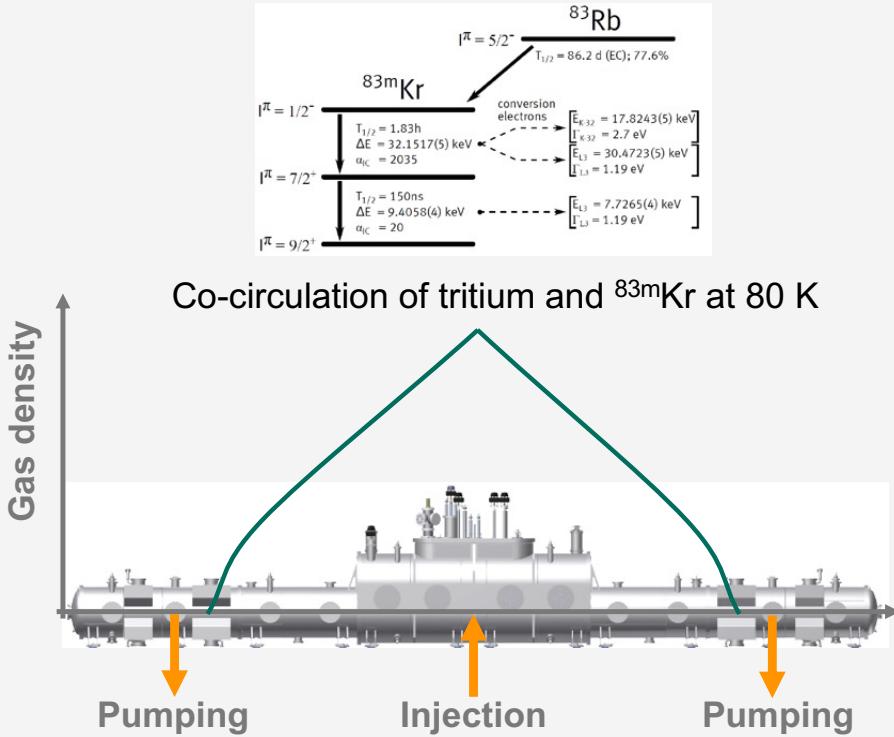


Systematics budget of analysis of 2nd campaign



arXiv:2105.08533 (subm. to Nature Physics)

Improving source-related systematics



Data of 2020 krypton run at 40% tritium column density used to constrain systematics in 2nd campaign [arXiv:2105.08533](https://arxiv.org/abs/2105.08533)

Since then: New operation mode with stable co-circulation at high column density at 80 K

Summer 2021: 10 GBq Krypton generator (activity x6) → further reduction of plasma systematics

Improving background

Main component of background (arXiv:2011.05107):

Highly excited (neutral) Rydberg atoms ionized in the volume of main spectrometer

Very low kinetic energy of background e^-

⇒ Volume dependent background rate

Reduce “downstream” volume of magnetic flux:
„shifted analyzing plane“ (SAP)

⇒ Factor 2 signal/background improvement

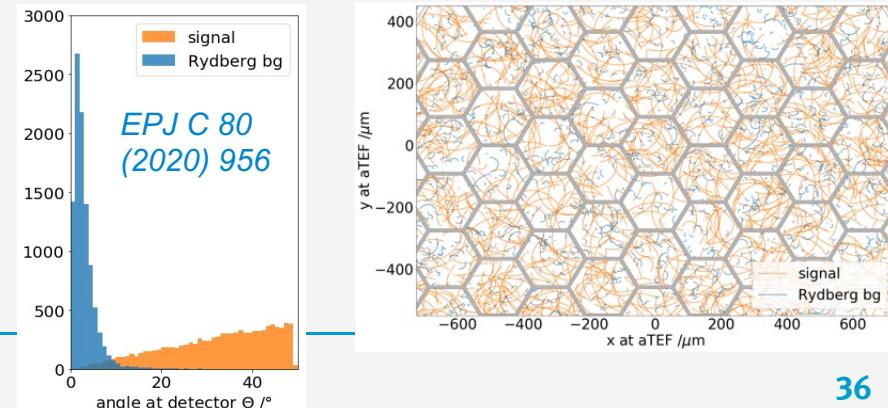
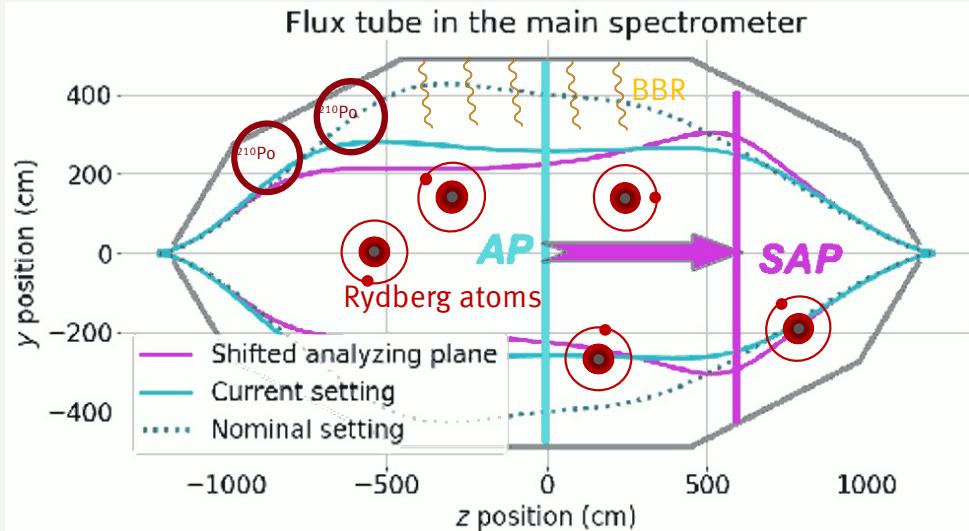
implemented in neutrino mass scans of 2020

Further reduction of background planned:

Make use of angular distribution of background electrons

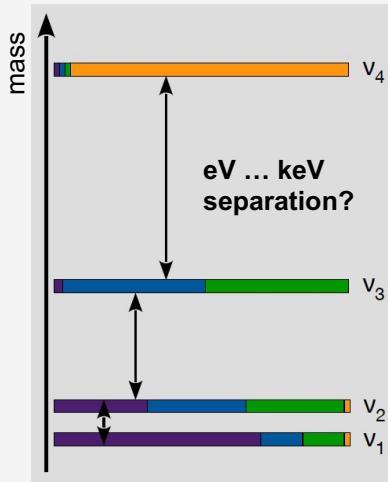
a very new idea: angular threshold filter

(“active transverse energy filter”, aTEF)

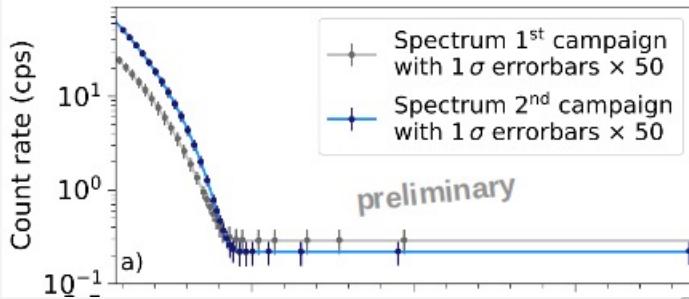


KATRIN „beyond the neutrino mass“

Is there a fourth (sterile) neutrino?



Neutrino mixing: “Kink” in normal β -spectrum (eV scale) or deep β -spectrum (keV scale)

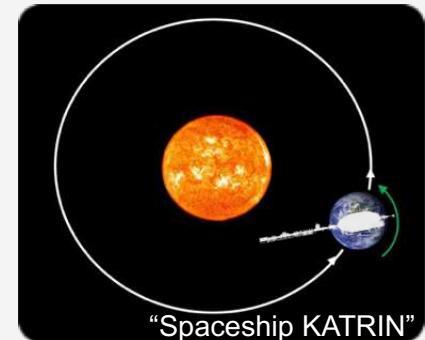


β -spectrum of high statistics and precision

Search for exotic weak interactions (spectrum shape)

Search for Lorentz invariance violation (sidereal modulation)

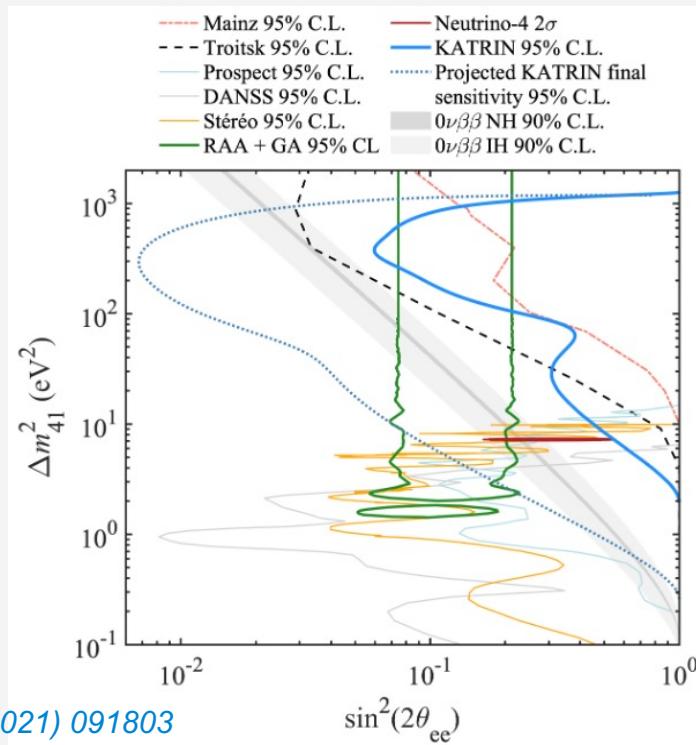
Constrain local overdensity of cosmic relic neutrinos (peak search)



Sterile neutrino searches in KATRIN

eV-scale sterile neutrinos:

new exclusion from the first science run
complementary probe of sterile ν



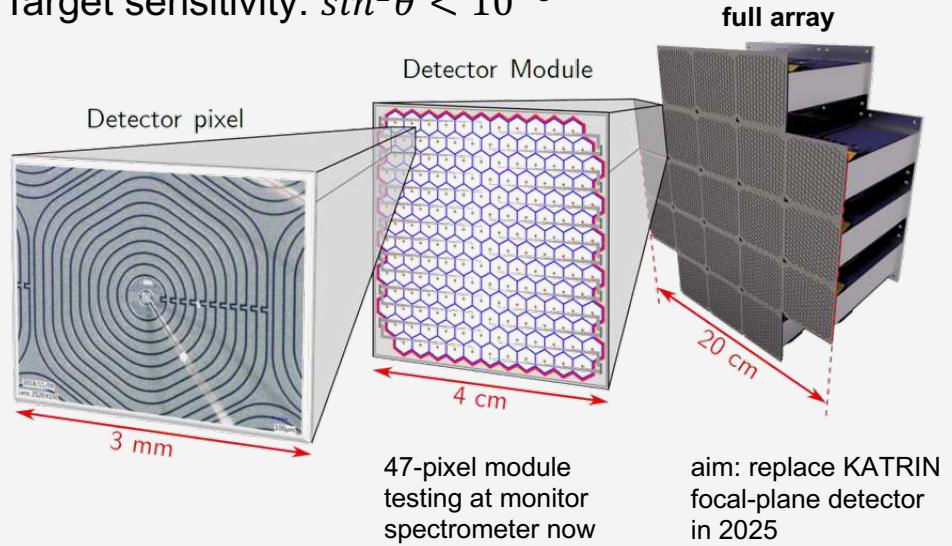
PRL 126 (2021) 091803

keV-scale sterile neutrinos: TRISTAN at KATRIN:

novel multi-pixel Silicon Drift Detector array
large count rates

excellent energy resolution

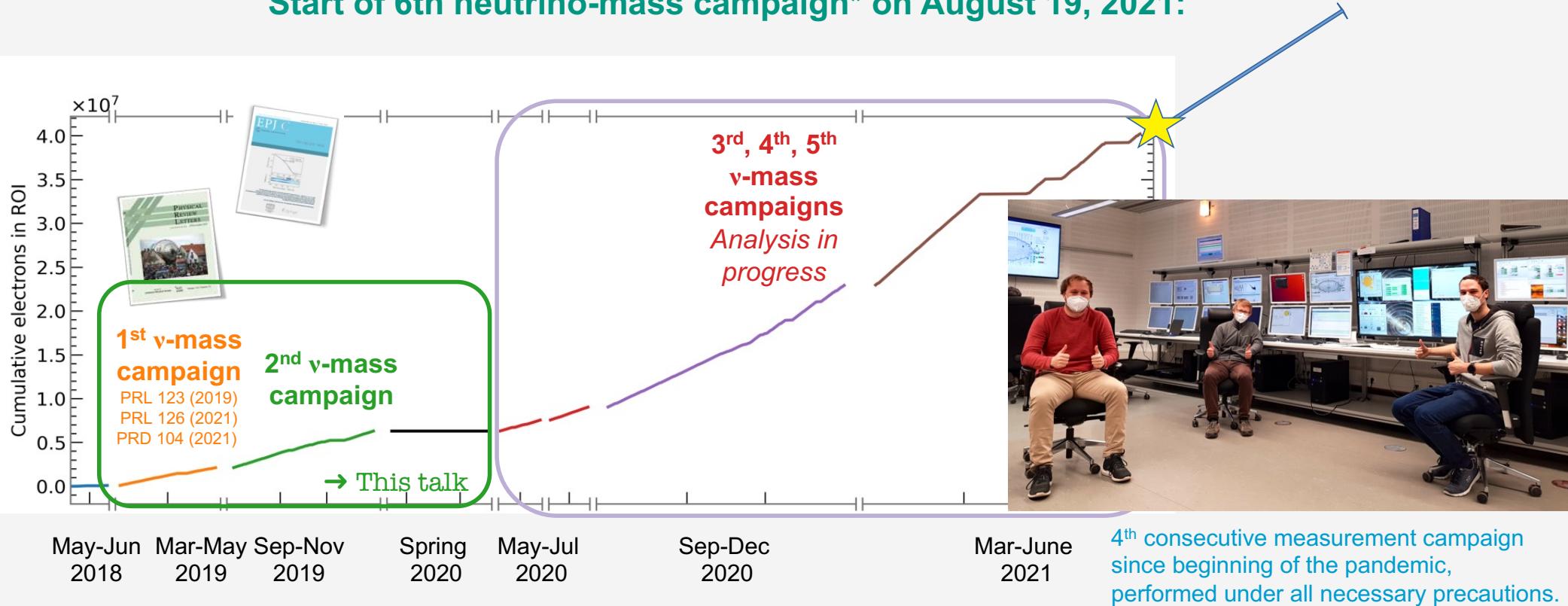
Target sensitivity: $\sin^2\theta < 10^{-6}$



J.Phys.G 46 (2019) 6, 065203, JINST 14 (2019) 11 P11013,
J. Phys. C. Ser. 1468 (2020) 012177

KATRIN data taking continues

Start of 6th neutrino-mass campaign* on August 19, 2021:



Conclusions

- ν physics remains super important for particle physics & beyond keeps being an important gateway to BSM physics
- **Hierarchy and δ_{CP}** are being resolved by LBL oscillation experiments with accelerator, reactor & atmospheric ν : now: T2K, NOvA, IceCube coming: JUNO, ORCA, Hyper Kamiokande, DUNE
- **Low energy puzzles:** $\nu_e, \bar{\nu}_e$ appearance, reactor neutrino anomaly, bump in reactor ν spectrum are being attacked by short baseline reactor experiments, SBN program, KATRIN new MicroBooNE results question MiniBooNE/LSND $\nu_e, \bar{\nu}_e$ appearance
- **CEvNS** has been established and will be an important tool for precision studies and BSM searches
- **Lepton number violation** will be searched for by $0\nu\beta\beta$ excellent final result by GERDA, LEGEND-200 will soon start to attack 10^{27} yr sensitivity CUPID, LEGEND-1000, nEXO: even more sensitive experiments with 10^{28} yr sensitivity
- **Direct neutrino mass experiments:** KATRIN presents 1st sub-eV limit, much more data, lower background rate and better systematics, goal: 0.2 eV sensitivity (new projects for the future to reach beyond KATRIN: ECHo, HOLMES, Project 8)

Conclusions

- ν physics remains super important for particle physics & beyond keeps being an important gateway to BSM physics
- Hierarchy and δ_{CP} are being resolved by LBL oscillation experiments with accelerator, reactor & atmospheric ν :
no
- **Many thanks for those who helped me with material & discussions
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S. Calvez, S.J. Dolan, G. Drexlin, J. Formaggio, A. Göttel, M. Gonchar, G. Gratta, P. Guzowski, C. Hati, T. Lasserre, A. Lokhov, M. Licciardi, M. Lindner, S. Mertens, R.C. Mandujano, S. Schönert, T. Tashiro, M. Uchita, K. Valerius
- **Thank you all for your attention**
- **Direct neutrino mass experiments:**
KATRIN presents 1st sub-eV limit
much more data, lower background rate and better systematics, goal: 0.2 eV sensitivity
(new projects for the future to reach beyond KATRIN: ECHo, HOLMES, Project 8)